

Reminder :
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of the perfect fluid
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Primitive equations
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RSW model
Primitive
equations (PE)
What we lose by
supposing
hydrostatics
Preliminary
conclusions.

Chapter 1: GFD models: reminder/derivations

V. Zeitlin

Cours GFD M2 OACOS

Plan

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GFD : space view

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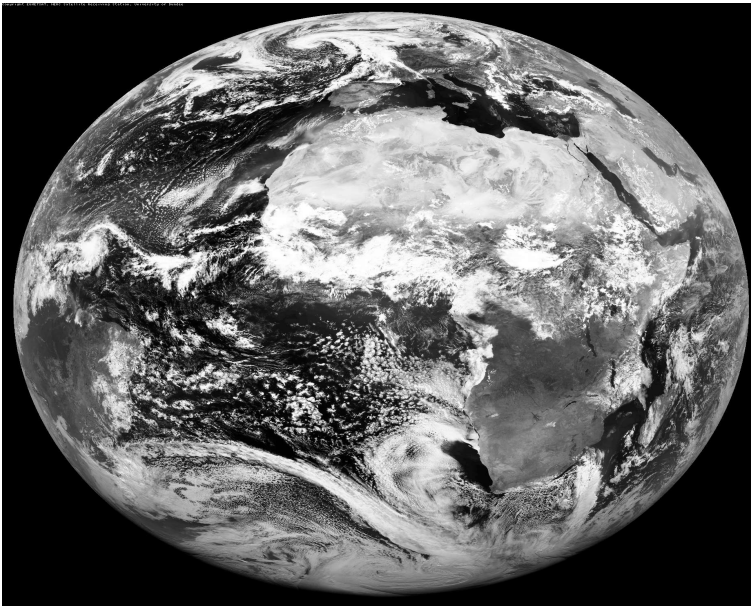
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Hydrodynamics in all its complexity plus :

- ▶ Rotating frame
- ▶ Thermal/stratification effects
- ▶ Spherical geometry (large- and meso-scales)
- ▶ Fluid in the complex domains (coasts, topography/bathymetry)
- ▶ Multi-phase fluid (water vapor, ice)

But !

These additional effects often allow to **simplify** the analysis

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Scales :

- ▶ Large : planetary 10^4 km
- ▶ Medium : atmosphere - synoptic, 10^3 km ; ocean - meso-scale $10 - 10^2$ km
- ▶ Small : atmosphere - meso-scale $1 - 10$ km ; ocean - sub-meso scale 1 km
- ▶ Very small : meters

Our prime interest : **large and medium scales.**

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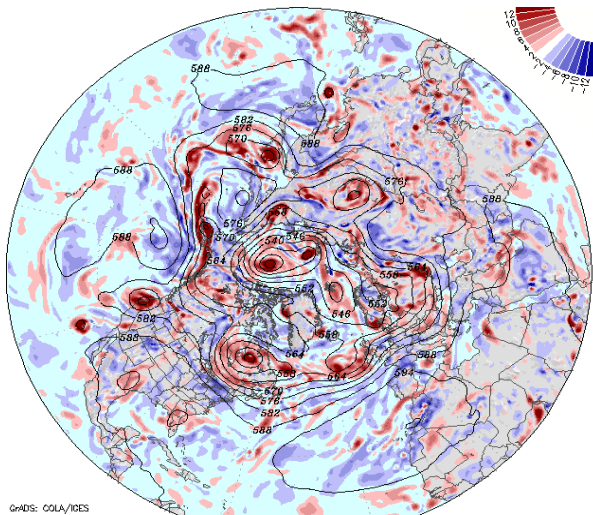
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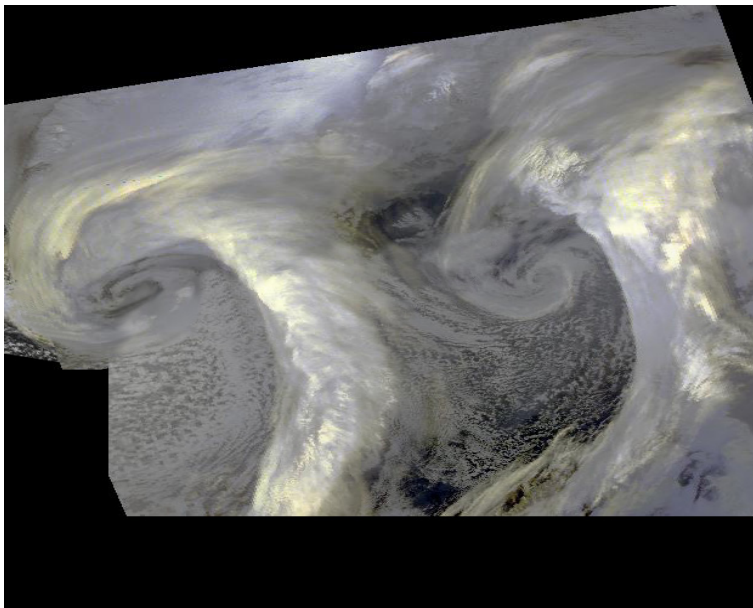


GrADS: GOLA/IGES

GFS Analysis: 00Z Wed 18 JUL 2012

500mb Geopotential Heights (dam), Vorticity ($10^{-4}/\text{sec}$)

Atmospheric vortices for real



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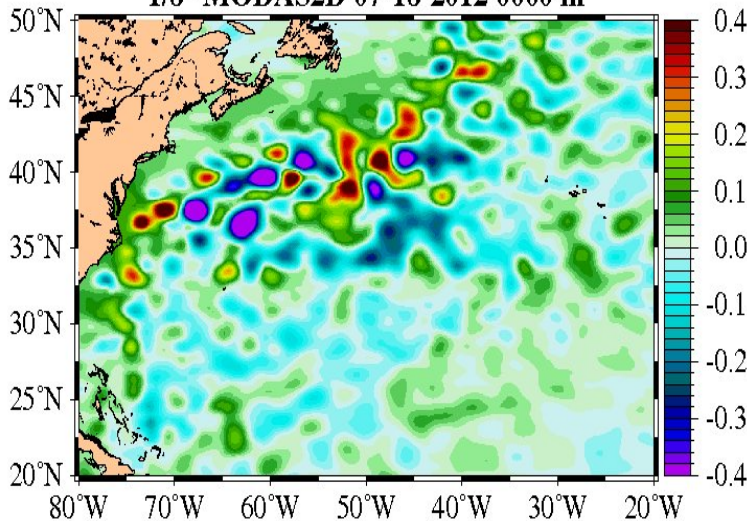
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Altimeter OI: Surface Height Deviation (m)
1/8° MODAS2D 07-18-2012 0000 m



Dynamical actors : waves, atmosphere

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Where the governing equations come from :

- ▶ **Mechanical system** \Rightarrow Newton's 2nd law \leftrightarrow momentum conservation.
- ▶ **Continuous medium** \Rightarrow local mass conservation
- ▶ **Thermodynamical system** \Rightarrow 1st and 2nd laws of thermodynamics, equation of state

Principal difficulty - **nonlinearity**

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Example of essentially nonlinear process : wave breaking



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Mathematical methods and tools

Related mathematical fields

- ▶ Linear algebra
- ▶ Partial differential equations
- ▶ Vector and tensor analyses
- ▶ Fourier analysis

Toolbox

- ▶ Method of small perturbations. Linearisation. Eigenproblems.
- ▶ Method of (time- and space-) averaging
- ▶ Asymptotic expansions, multi-scale analysis

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Fluid dynamics according to Lagrange :

Description in terms of instantaneous positions of **fluid parcels** $\vec{X}(\vec{x}, t)$, along their trajectories, where \vec{x} are initial positions (Lagrangian labels).

Newton's 2nd law :

$$\rho(\vec{X}, t) \frac{d^2 \vec{X}}{dt^2} = -\vec{\nabla} P(\vec{X}, t). \quad (1)$$

Continuity equation :

$$\rho_i(x) d^3 \vec{x} = \rho(\vec{X}, t) d^3 \vec{X}, \leftrightarrow \rho_i(x) = \rho(\vec{X}, t) \mathcal{J} \quad (2)$$

where ρ_i is initial distribution of density of the fluid,
 $\mathcal{J} = \frac{\partial(X, Y, Z)}{\partial(x, y, z)}$ is the Jacobi determinant d(Jacobian). Fluid
velocity : $\vec{v}(\vec{X}, t) = \frac{d\vec{X}}{dt} \equiv \dot{\vec{X}}$.

Fluid dynamics according to Euler :

Description in terms of instantaneous values of the velocity, density and pressure fields at the **fixed point** of space :

$\vec{v}(\vec{x}, t)$, $\rho(\vec{x}, t)$, $P(\vec{x}, t)$. Duality : $\vec{X} \leftrightarrow \vec{x}$

Newton's 2nd law :

$$\rho \left(\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} \right) = -\vec{\nabla} P. \quad (3)$$

Continuity equation :

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v}) = 0. \quad (4)$$

Lagrangian derivative :

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \vec{v} \cdot \vec{\nabla}. \quad (5)$$

Proposition

Lagrangian and Eulerian continuity equations are equivalent

Démonstration.

$$\frac{d}{dt}(\rho \mathcal{J}) = \frac{d\rho}{dt} \mathcal{J} + \rho \frac{d\mathcal{J}}{dt} = \frac{d\rho_i}{dt} = 0, \quad (6)$$

$$\begin{aligned} \frac{d\mathcal{J}}{dt} &= \frac{\partial(\dot{X}, Y, Z)}{\partial(x, y, z)} + \frac{\partial(X, \dot{Y}, Z)}{\partial(x, y, z)} + \frac{\partial(X, Y, \dot{Z})}{\partial(x, y, z)} \\ &= \left(\frac{\partial(\dot{X}, Y, Z)}{\partial(X, Y, Z)} + \dots \right) \mathcal{J} = \left(\frac{\partial \dot{X}}{\partial X} + \frac{\partial \dot{Y}}{\partial Y} + \frac{\partial \dot{Z}}{\partial Z} \right) \mathcal{J} \Rightarrow \end{aligned}$$

$$\frac{d\rho}{dt} + \rho \vec{\nabla} \cdot \vec{v} = 0 \Leftrightarrow \frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v}) = 0. \quad (7)$$

□

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Closure of the system : equation of state

General equation of state

$$P = P(\rho, s), \quad (8)$$

wher s - entropy per unit mass ;

► **Barotropic** fluid :

$$P = P(\rho) \leftrightarrow s = \text{const}, \quad (9)$$

► **Baroclinic** fluid :

$$P = P(\rho, s), \Rightarrow \quad (10)$$

Equation for s necessary. **Perfect fluid** :

$$\frac{ds}{dt} = \frac{\partial s}{\partial t} + \vec{v} \cdot \vec{\nabla} s = 0. \quad (11)$$

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Particular case of the barotropic fluid - incompressible fluid :

Volume conservation :

$$\mathcal{J} = 1 \Leftrightarrow \vec{\nabla} \cdot \vec{v} = 0 \Rightarrow . \quad (12)$$

pressure is not independent variable.

1. If in addition, $\rho = \text{const}$:

$$\vec{\nabla} \cdot (\vec{v} \cdot \vec{\nabla} \vec{v}) = -\frac{1}{\rho} \vec{\nabla}^2 P. \quad (13)$$

2. Otherwise

$$\frac{d\rho}{dt} = \frac{\partial \rho}{\partial t} + \vec{v} \cdot \vec{\nabla} \rho = 0. \quad (14)$$

et

$$\vec{\nabla} \cdot (\vec{v} \cdot \vec{\nabla} \vec{v}) = -\vec{\nabla} \cdot \left(\frac{\vec{\nabla} P}{\rho} \right). \quad (15)$$

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Thermodynamics : reminder

1st principle, "dry" thermodynamics

$$\delta\epsilon = T\delta s - P\delta v, \quad (16)$$

where ϵ - internal energy per unit mass, $v = \frac{1}{\rho}$ - volume per unit mass.

Enthalpy per unit mass : $h = \epsilon + Pv$:

$$\delta h = T\delta s + v\delta P. \quad (17)$$

Energy density of the fluid :

$$e = \frac{\rho \vec{v}^2}{2} + \rho\epsilon. \quad (18)$$

Local conservation of energy :

$$\frac{\partial e}{\partial t} + \vec{\nabla} \cdot \left[\rho \vec{v} \left(\frac{\vec{v}^2}{2} + h \right) \right] = 0. \quad (19)$$

Kelvin theorem

Circulation :

$$\gamma = \int_{\Gamma} \vec{v} \cdot d\vec{l} = \int_{S_{\Gamma}} (\vec{\nabla} \wedge \vec{v}) \cdot d\vec{l}, \quad (21)$$

where Γ - arbitrary contour, S_{Γ} - surface with the boundary Γ .

Kelvin theorem

- Barotropic fluid

$$\frac{d\gamma}{dt} = 0, \quad (22)$$

- Baroclinic fluid

$$\frac{d\gamma}{dt} = - \int_{\Gamma} \frac{\vec{\nabla} P}{\rho} \cdot d\vec{l}, \quad (23)$$

Kelvin theorem

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Exercise

- ▶ Prove energy conservation and Kelvin theorem for the barotropic fluid
- ▶ Same for the baroclinic fluid
- ▶ Write down, with demonstration, the Euler equations for the incompressible fluid in cylindrical coordinates

Dissipative phenomena : molecular fluxes

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Effects of dissipation : correction of the macroscopic fluxes
of :

- ▶ momentum
- ▶ mass
- ▶ internal energy (heat)

by the corresponding molecular fluxes, calculated from the
flux - gradient relations :

$$\vec{f}_A = -k_A \vec{\nabla} A, \quad (24)$$

A - a thermodynamical variable, \vec{f}_A - corresponding molecular
flux.

Viscosity

Tensor notation

$$\vec{x} \rightarrow x_i, \quad \vec{v} \rightarrow v_i, \quad \vec{\nabla} \rightarrow \partial_i, \quad i = 1, 2, 3. \quad (25)$$

Einstein's convention : repeating indices - summation from 1 to 3.

Conservation of the momentum :

$$\partial_t(\rho v_i) + \partial_k \pi_{ik} = 0, \quad \pi_{ik} = \rho v_i v_k + P \delta_{ik}, \quad \delta_{ik} = \text{diag}(1, 1, 1). \quad (26)$$

Viscous tensions - (density of) the molecular flux of the momentum :

$$\sigma_{ik} = \nu(\partial_i v_k + \partial_k v_i) \Rightarrow \partial_t(\rho v_i) + \partial_k(\pi_{ik} - \sigma_{ik}) = 0, \quad (27)$$

Incompressible case : Navier -Stokes (NS) equation

$$\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} = -\frac{\vec{\nabla} P}{\rho} + \nu \vec{\nabla}^2 \vec{v}, \quad \vec{\nabla} \cdot \vec{v} = 0. \quad (28)$$

Reynolds' number

Dimensionless form of the NS equation :

$$\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} = -\frac{\vec{\nabla} P}{\rho} + \frac{1}{Re} \vec{\nabla}^2 \vec{v}, \quad (29)$$

$Re = UL/\nu$, U , L -typical velocity- and length-scales.

Remarque : typical Re for synoptic motions $\rightarrow \infty$

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Diffusivity, thermal conductivity

Molecular fluxes of mass and heat :

$$-D\vec{\nabla}\rho, \quad -\kappa\vec{\nabla}T \quad (30)$$

Corrected continuity equation :

$$\frac{\partial\rho}{\partial t} + \vec{\nabla} \cdot (\rho\vec{v}) = D\vec{\nabla}^2\rho. \quad (31)$$

Equation of heat/temperature

$$\frac{\partial T}{\partial t} + \vec{v} \cdot \vec{\nabla}T = \chi\vec{\nabla}^2T. \quad (32)$$

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Motion in the rotating frame

Material point in the rotating frame :

$$m \frac{d\vec{v}}{dt} + 2m\vec{\Omega} \wedge \vec{v} + m\vec{\Omega} \wedge (\vec{\Omega} \wedge \vec{r}) = 0, \quad \vec{v} = \frac{d\vec{r}}{dt} \quad (33)$$

Euler equations in the rotating frame in the presence of gravity :

$$\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} + 2\vec{\Omega} \wedge \vec{v} - \vec{g}^* = -\frac{\vec{\nabla} P}{\rho} \quad (34)$$

Effective gravity :

$$\vec{g}^* = \vec{g} + m\vec{\Omega} \wedge (\vec{\Omega} \wedge \vec{r}) \quad (35)$$

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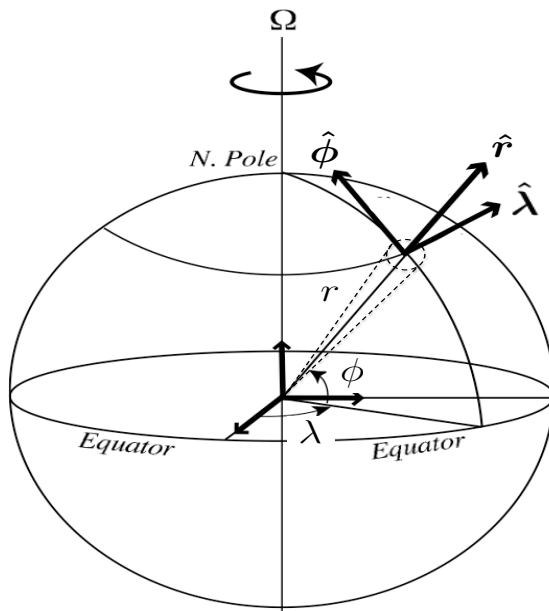
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$$\begin{aligned} \frac{dv_r}{dt} - \frac{v_\lambda^2 + v_\phi^2}{r} - 2\Omega \cos \phi v_\lambda + g^* &= -\frac{1}{\rho} \partial_r P, \\ \frac{dv_\lambda}{dt} + \frac{v_r v_\lambda - v_\phi v_\lambda \tan \phi}{r} + 2\Omega (-\sin \phi v_\phi + \cos \phi v_r) &= -\frac{1}{\rho r \cos \phi} \partial_\lambda P, \\ \frac{dv_\phi}{dt} + \frac{v_r v_\phi + v_\lambda^2 \tan \phi}{r} + 2\Omega \sin \phi v_\lambda &= -\frac{1}{\rho r} \partial_\phi P, \\ \frac{d\rho}{dt} + \rho \left[\frac{1}{r^2} \frac{\partial(r^2 v_r)}{\partial r} + \frac{1}{r \cos \phi} \left(\frac{\partial(\cos \phi v_\phi)}{\partial \phi} + \frac{\partial v_\lambda}{\partial \lambda} \right) \right], \\ \frac{d}{dt} &= \frac{\partial}{\partial t} + v_r \partial_r + \frac{v_\phi}{r} \partial_\phi + \frac{v_\lambda}{r \cos \phi} \partial_\lambda \end{aligned}$$

Traditional approx. : green + red \rightarrow out, $r \rightarrow R = \text{const}$

Non-traditional approx : green \rightarrow out.

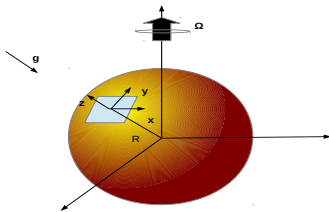
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$$\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} + f \hat{z} \wedge \vec{v} + \vec{g} = -\frac{\vec{\nabla} P}{\rho}$$

f - plane : $f = \text{const}$; β - plane : $f = f + \beta y$; f - Coriolis
parameter : $f = 2\Omega \sin \phi$

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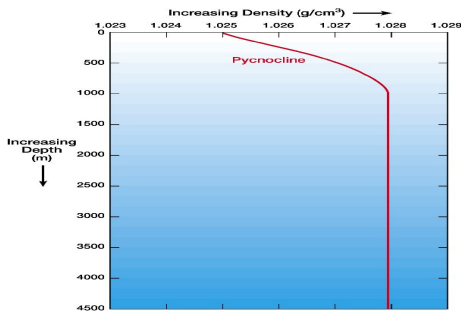
Exercise

Deduce Euler and continuity equations in spherical coordinates.

Determine conditions of validity of the tangent plane approximation.

Ocean : observations and approximations

Typical density profile :



$$\rho(\vec{x}, t) = \rho_0 + \rho_s(z) + \sigma(x, y, z; t), \quad \rho_0 \gg \rho_s \gg \sigma. \quad (36)$$

Hydrostatics

$$g\rho + \partial_z P = 0, \Rightarrow P = P_0 + P_s(z) + \pi(x, y, z; t), \quad (37)$$

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Equations of motion

$$\vec{\nabla} \cdot \vec{v} = 0, \quad \vec{v} = \vec{v}_h + \hat{z}w. \quad (38)$$

Euler equations :

$$\frac{\partial \vec{v}_h}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v}_h + f \hat{z} \wedge \vec{v}_h = -\frac{\vec{\nabla}_h \pi}{\rho} \approx -\vec{\nabla}_h \phi. \quad (39)$$

$\phi = \frac{\pi}{\rho_0}$ - **geopotential**.

Continuity equation :

$$\partial_t \rho + \vec{v} \cdot \vec{\nabla} \rho = 0. \quad (40)$$

Boundary conditions : Rigid lid and flat bottom :

$$w|_{z=0} = w|_{z=H} = 0 \quad (41)$$

Non-trivial bathymetry : $w|_{z=b} = \frac{db}{dt}$

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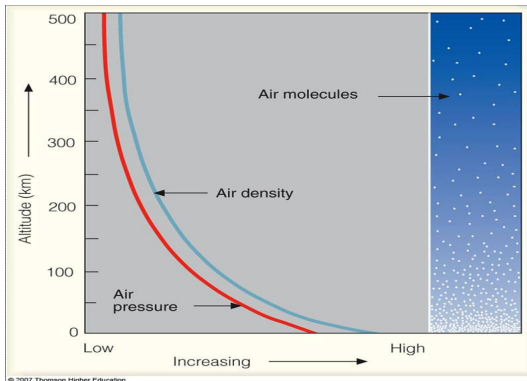
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Altitude \leftrightarrow Pressure \Rightarrow vertical coordinate.

Thermodynamics of the dry atmosphere

Equation of state - ideal gas :

$$P = \rho RT, \quad c_{P,V} = T \left(\frac{\partial s}{\partial T} \right)_{P,V} = \text{const}, \quad c_p - c_v = R. \quad (42)$$

Entropy :

$$s = c_p \ln T - R \ln P + \text{const}. \quad (43)$$

Adiabatic process :

$$s = \text{const} \Rightarrow c_p \frac{dT}{T} - R \frac{dP}{P} = 0, \Rightarrow T = T_s \left(\frac{P}{P_s} \right)^{\frac{R}{c_p}}. \quad (44)$$

Potential temperature :

$$\theta = T \left(\frac{P_s}{P} \right)^{\frac{R}{c_p}}, \quad s = c_p \ln \theta + \text{const}. \quad (45)$$

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Geopotential

$\delta\phi = g\delta z$, où $z = z(p)$ via hydrostatics,
 z - thermodynamical variable.

Hydrostatics

$$\delta\phi = -\frac{RT}{P}\delta P \Rightarrow \quad (46)$$

$$\frac{\partial\phi}{\partial p} = -\frac{RT}{P} = -\frac{1}{\rho}. \quad (47)$$

Elimination of ρ

"Triangular" relation :

$$\left(\frac{\partial P}{\partial x}\right)_z \left(\frac{\partial x}{\partial z}\right)_P \left(\frac{\partial z}{\partial P}\right)_x = -1 \Rightarrow \quad (48)$$

$$\left(\frac{\partial P}{\partial x}\right)_z = - \left(\frac{\partial P}{\partial z}\right)_x \left(\frac{\partial z}{\partial x}\right)_P = \rho \left(\frac{\partial \phi}{\partial x}\right)_P. \quad (49)$$

Incompressibility in pressure coordinates

Lagrangian volume element in pressure coordinates :

$$\rho dx dy dz = -\frac{1}{g} dx dy dP \quad (50)$$

Mass conservation \Rightarrow Volume conservation in P .

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$$\operatorname{div}(\vec{v}) = \vec{\nabla}_h \cdot \vec{v}_h + \partial_p \omega = 0, \quad \omega = \frac{dP}{dt}. \quad (51)$$

$$\frac{\partial \vec{v}_h}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v}_h + f \hat{z} \wedge \vec{v}_h = -\vec{\nabla}_h \phi. \quad (52)$$

$$\partial_t \theta + \vec{v} \cdot \vec{\nabla} \theta = 0. \quad (53)$$

$$\frac{\partial \phi}{\partial p} = -\frac{RT}{P} = -\frac{R}{P} \left(\frac{P}{P_s} \right)^{\frac{R}{c_p}} \theta. \quad (54)$$

"Pseudo-height" coordinate

New vertical coordinate :

$$\bar{z} = z_0 \left(1 - \left(\frac{P}{P_s} \right)^{\frac{R}{c_p}} \right) \equiv z_0 \left(1 - \left(\frac{P}{P_s} \right)^{\frac{\gamma-1}{\gamma}} \right), \quad (55)$$

$$z_0 = \frac{\gamma}{\gamma - 1} \frac{P_s}{g \rho_s} \approx 28 \text{ km}. \quad (56)$$

Pseudo- density :

$$r : \quad r d\bar{z} = \rho dz = -\frac{1}{g} dP. \quad (57)$$

Mass conservation :

$$dx dy dP = -g r(\bar{z}) dx dy d\bar{z} \Rightarrow \quad (58)$$

$$r \left(\vec{\nabla}_h \cdot \vec{v}_h + \frac{\partial \bar{w}}{\partial \bar{z}} \right) + \bar{w} \frac{\partial r}{\partial \bar{z}} = 0, \quad \vec{v} = (\vec{v}_h, \bar{w} = \dot{\bar{z}}). \quad (59)$$

Approximation $\bar{z} \ll z_0$:

$$\vec{\nabla}_h \cdot \vec{v}_h + \frac{\partial \bar{w}}{\partial \bar{z}} = -\frac{\bar{w}}{r} \frac{\partial r}{\partial \bar{z}} = \frac{\bar{w}}{(\gamma - 1) z_0 \left(1 - \frac{\bar{z}}{z_0} \right)} \approx 0. \quad (60)$$

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$$\frac{\partial \vec{v}_h}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v}_h + f \hat{z} \wedge \vec{v}_h = -\vec{\nabla}_h \phi, \quad (61)$$

$$-g \frac{\theta}{\theta_0} + \frac{\partial \phi}{\partial \bar{z}} = 0, \quad (62)$$

$$\frac{\partial \theta}{\partial t} + \vec{v} \cdot \vec{\nabla} \theta = 0; \quad \vec{\nabla} \cdot \vec{v} = 0. \quad (63)$$

Identical to oceanic equations with $\sigma \rightarrow -\theta$.

Isentropic coordinates

Montgomery potential

$$\psi = \phi - \frac{\theta}{\theta_0} g \bar{z}, \Rightarrow \quad (64)$$

$$\begin{aligned} d\psi &= d\phi - g \frac{\theta}{\theta_0} d\bar{z} - g d\frac{\theta}{\theta_0} \bar{z} \\ &= \vec{\nabla}_h \phi \cdot d\vec{x}_h + \partial_{\bar{z}} \phi d\bar{z} - \frac{\theta}{\theta_0} g d\bar{z} - g \bar{z} d\frac{\theta}{\theta_0} \\ &= \vec{\nabla}_h \phi \cdot d\vec{x}_h - g \bar{z} d\frac{\theta}{\theta_0} \end{aligned} \quad (65)$$

Therefore :

$$\left(\vec{\nabla}_h \psi \right)_\theta = \left(\vec{\nabla}_h \phi \right)_{\bar{z}} ; \partial_\theta \psi = -g \bar{z} / \theta_0 \quad (66)$$

and \bar{z} is a new dependent variable.

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Velocity and continuity equation in isentropic coordinates

Velocity

$$\vec{v} \rightarrow \left(\vec{v}_h, \tilde{w} = \frac{d\theta}{dt} \right) \Rightarrow \quad (67)$$

$\tilde{w} \equiv 0$ for adiabatic processes

Mass conservation

$$dx dy d\bar{z} = \frac{\partial \bar{z}}{\partial \theta} dx dy d\theta = \text{const} \rightarrow \quad (68)$$

$$\partial_t \left(\frac{\partial \bar{z}}{\partial \theta} \right) + \vec{\nabla}_h \cdot \left(\frac{\partial \bar{z}}{\partial \theta} \vec{v}_h \right) = 0. \quad (69)$$

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$$\frac{\partial \vec{v}_h}{\partial t} + \vec{v}_h \cdot \vec{\nabla}_h \vec{v}_h + f \hat{z} \wedge \vec{v}_h = -\vec{\nabla}_h \psi, \quad (70)$$

$$+ \frac{g\bar{z}}{\theta_0} + \frac{\partial \psi}{\partial \theta} = 0, \quad (71)$$

$$\partial_t \left(\frac{\partial \bar{z}}{\partial \theta} \right) + \vec{\nabla}_h \cdot \left(\frac{\partial \bar{z}}{\partial \theta} \vec{v}_h \right) = 0. \quad (72)$$

Rewriting Euler equations in rotating frame in form of conservation laws and vertical integration

Equations of horizontal motion

$$\partial_t(\rho u) + \partial_x(\rho u^2) + \partial_y(\rho v u) + \partial_z(\rho w u) - f \rho v = -\partial_x p, \quad (73)$$

$$\partial_t(\rho v) + \partial_x(\rho u v) + \partial_y(\rho v^2) + \partial_z(\rho w v) + f \rho u = -\partial_y p, \quad (74)$$

Vertical integration

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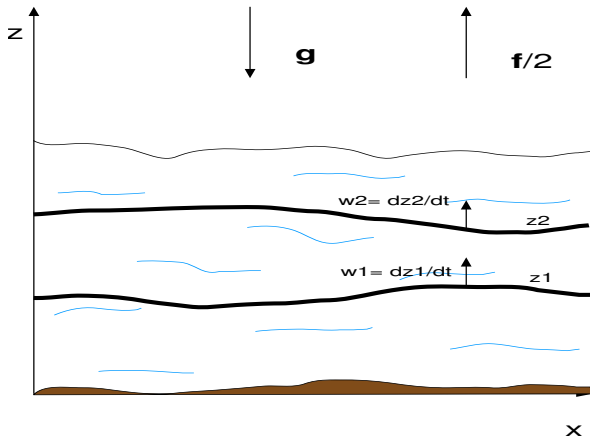
Integration between two material surfaces $z_{1,2}$.

By definition :

$$w|_{z_i} = \frac{dz_i}{dt} = \partial_t z_i + u \partial_x z_i + v \partial_y z_i, \quad i = 1, 2. \quad (75)$$

Leibnitz formula :

$$\int_{z_1}^{z_2} dz \partial_x F = \partial_x \int_{z_1}^{z_2} dz F - \partial_x z_2 F|_{z_2} + \partial_x z_1 F|_{z_1} \quad (76)$$



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Integrated equations

Using (75) and (76) we obtain :

$$\begin{aligned} & \partial_t \int_{z_1}^{z_2} dz \rho u + \partial_x \int_{z_1}^{z_2} dz \rho u^2 + \partial_y \int_{z_1}^{z_2} dz \rho uv \\ & - f \int_{z_1}^{z_2} dz \rho v = -\partial_x \int_{z_1}^{z_2} dz p - \partial_x z_1 p|_{z_1} + \partial_x z_2 p|_{z_2} . \end{aligned}$$

$$\begin{aligned} & \partial_t \int_{z_1}^{z_2} dz \rho v + \partial_x \int_{z_1}^{z_2} dz \rho uv + \partial_y \int_{z_1}^{z_2} dz \rho v^2 \\ & + f \int_{z_1}^{z_2} dz \rho u = -\partial_y \int_{z_1}^{z_2} dz p - \partial_y z_1 p|_{z_1} + \partial_y z_2 p|_{z_2} . \end{aligned}$$

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$$\partial_t \int_{z_1}^{z_2} dz \rho + \partial_x \int_{z_1}^{z_2} dz \rho u + \partial_y \int_{z_1}^{z_2} dz \rho v = 0. \quad (77)$$

Integrated density :

$$\mu = \int_{z_1}^{z_2} dz \rho = -\frac{1}{g} (p|_{z_2} - p|_{z_1}), \quad (78)$$

Density-weighted vertical average :

$$\langle F \rangle = \frac{1}{\mu} \int_{z_1}^{z_2} dz \rho F. \quad (79)$$

Equations for the averages :

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$$\begin{aligned} \partial_t (\mu \langle u \rangle) &+ \partial_x (\mu \langle u^2 \rangle) + \partial_y (\mu \langle uv \rangle) - f \mu \langle v \rangle = \\ &- \partial_x \int_{z_1}^{z_2} dz p - \partial_x z_1 p|_{z_1} + \partial_x z_2 p|_{z_2}, \quad (80) \end{aligned}$$

$$\begin{aligned} \partial_t (\mu \langle v \rangle) &+ \partial_x (\mu \langle uv \rangle) + \partial_y (\mu \langle v^2 \rangle) + f \mu \langle u \rangle = \\ &- \partial_y \int_{z_1}^{z_2} dz p - \partial_y z_1 p|_{z_1} + \partial_y z_2 p|_{z_2}, \quad (81) \end{aligned}$$

$$\partial_t \mu + \partial_x (\mu \langle u \rangle) + \partial_y (\mu \langle v \rangle) = 0. \quad (82)$$

Mean-field approximation :

Expression for the pressure

Pressure inside the layer (z_1, z_2) in terms of pressure at the lower surface and vertical position :

$$p(x, y, z, t) = -g \int_{z_1}^z dz' \rho(x, y, z', t) + p|_{z_1}. \quad (83)$$

Closure hypothesis :

Weak variations in the vertical, **correlations decoupled** :

$$\langle uv \rangle \approx \langle u \rangle \langle v \rangle, \quad \langle u^2 \rangle \approx \langle u \rangle \langle u \rangle, \quad \langle v^2 \rangle \approx \langle v \rangle \langle v \rangle. \quad (84)$$

Approximate equations

Mean density :

Mean density $\bar{\rho}$:

$$\bar{\rho} = \frac{1}{(z_2 - z_1)} \int_{z_1}^{z_2} dz \rho, \quad \mu = \bar{\rho}(z_2 - z_1). \quad (85)$$

Pressure in terms of $\bar{\rho}$:

$$p(x, y, z, t) \approx -g\bar{\rho}(z - z_1) + p|_{z_1}. \quad (86)$$

Hypothesis : $\bar{\rho} = \text{const}$ in what follows.

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Omitting the brackets we obtain for the averages from (80), (81), (84), (86), with the help of (82), (85) :

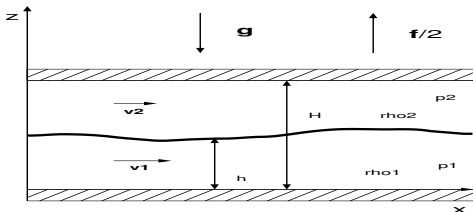
$$\begin{aligned}
 \bar{\rho}(z_2 - z_1)(\partial_t \mathbf{v}_h + \mathbf{v}_h \cdot \nabla_h \mathbf{v}_h + f \hat{\mathbf{z}} \wedge \mathbf{v}_h) = \\
 -\nabla_h \left(-g \bar{\rho} \frac{(z_2 - z_1)^2}{2} + (z_2 - z_1) p|_{z_1} \right) \\
 -\nabla_h z_1 p|_{z_1} + \nabla_h z_2 p|_{z_2} .
 \end{aligned} \tag{87}$$

Any variable in this equation is a function only of horizontal coordinates and time. Alternative $v : \vec{v}_h \equiv \mathbf{v}_h$.

Rotating shallow water (RSW), 2 layers

2- layer configuration, rigid lid

Application of general equations (87) to the fluid between the bottom $z_1 = 0$ and the top $z_3 = H$ planes. Choose a material surface $z = z_2(x, y, t) \equiv h(x, y, t)$ in the fluid interior, $\vec{\nabla}_h \rightarrow \vec{\nabla}$, $\vec{v}_h \rightarrow \mathbf{v}$. Vertical boundaries - material surfaces. Generalisation to non-trivial topography : $z_1 = b(x, y)$.



Equations of motion

$\mathbf{v}_{1(2)}, \bar{\rho}_{1(2)}$ - velocity and density in the inferior (superior) layer.

$$\partial_t \mathbf{v}_2 + \mathbf{v}_2 \cdot \nabla \mathbf{v}_2 + f \hat{\mathbf{z}} \wedge \mathbf{v}_2 = -\frac{1}{\bar{\rho}_2} \nabla p|_H \quad (88)$$

$$\partial_t \mathbf{v}_1 + \mathbf{v}_1 \cdot \nabla \mathbf{v}_1 + f \hat{\mathbf{z}} \wedge \mathbf{v}_1 = -\frac{1}{\bar{\rho}_1} \nabla p|_H - g \frac{\bar{\rho}_1 - \bar{\rho}_2}{\bar{\rho}_1} \nabla h, \quad (89)$$

$$\partial_t h + \nabla \cdot (\mathbf{v}_1 h) = 0, \quad (90)$$

$$\partial_t (H - h) + \nabla \cdot (\mathbf{v}_2 (H - h)) = 0, \quad (91)$$

Re-interpretation of the equations

2-dimensional Euler equations in each layer with **dynamical boundary condition at the interface** :

$$p_1 = (\bar{\rho}_1 - \bar{\rho}_2)gh + p_2, \quad (92)$$

Reduced gravity

Remark : g enter equations uniquely in combination $g \frac{\bar{\rho}_1 - \bar{\rho}_2}{\bar{\rho}_1}$ - **reduced gravity**

Exercise

Deduce the equations (88) - (91).

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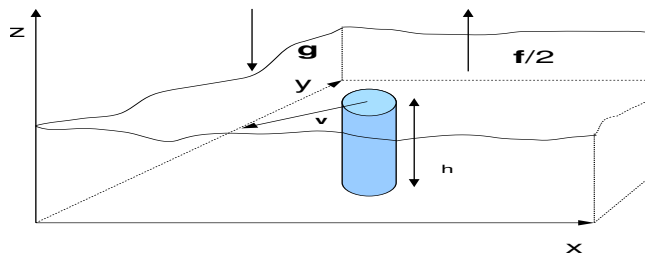
1-layer rotating shallow water model (Saint-Venant)

In the limit $\bar{\rho}_2 \rightarrow 0$:

$$\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v} + f \hat{\mathbf{z}} \wedge \mathbf{v} + g \nabla h = 0, \quad (93)$$

$$\partial_t h + \nabla \cdot (\mathbf{v} h) = 0, \quad (94)$$

In the presence of non-trivial topography $h \rightarrow h - b(x, y)$ in the second equation.



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Energy

By construction, equations (93), (94) express the local conservation of the horizontal momentum and mass. Energy density :

$$e = h \frac{\mathbf{v}^2}{2} + g \frac{h^2}{2} \quad (95)$$

obeys the conservation equation :

$$\partial_t e + \nabla \cdot \left(\mathbf{v} h \left(\frac{\mathbf{v}^2}{2} + gh \right) \right) = 0, \quad (96)$$

and total energy, $E = \int dx dy e$, is constant for an isolated system.

Potential vorticity, RSW model

Specific Lagrangian conservation law : potential vorticity q (PV), constructed from the **relative vorticity** (vertical component) $\zeta = v_x - u_y$, Coriolis le parametre f , and fluid depth h .

$$q = \frac{\zeta + f}{h}. \quad (97)$$

here $\zeta + f$ - **absolute vorticity** , f - **planetary vorticity**.

Lagrangian conservation :

$$\frac{dq}{dt} \equiv (\partial_t + \mathbf{v} \cdot \nabla) q = 0, \quad (98)$$

is obtained by combining equations of vorticity :

$$\frac{d(\zeta + f)}{dt} + (\zeta + f) \nabla \cdot \mathbf{v} = 0, \quad (99)$$

and continuity

$$\frac{dh}{dt} + h \nabla \cdot \mathbf{v} = 0 : \quad (100)$$

$$\frac{d}{dt} \frac{\zeta + f}{h} = \frac{1}{h} \frac{d}{dt} (\zeta + f) - \frac{\zeta + f}{h^2} \frac{d}{dt} h = 0, \quad (101)$$

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Eulerian expression :

Conservation of PV leads to independence of time of any integral :

$$\int dx dy h \mathcal{F}(q), \quad (102)$$

over the whole flow, with \mathcal{F} - arbitrary function.

Qualitative image of the RSW dynamics :

Two-dimensional motion of the fluid columns of variable depth, each preserving its potential vorticity.

Spectrum of the small perturbations - RSW model

Geophysical
Fluid Dynamics 1

V Zeitlin - GFD

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Linearised equations :

Perturbations about state of rest : $\mathbf{v} = 0$, $h = H_0 = \text{const.}$

Linearised equations in the approximation $f = f_0 = \text{const.}$:

$$\begin{aligned}u_t - fv + g\eta_x &= 0, \\v_t + fu + g\eta_y &= 0, \\ \eta_t + H_0(u_x + v_y) &= 0,\end{aligned}\tag{103}$$

where u , v - 2 components of the velocity perturbation, η - perturbation of the interface.

Method of Fourier

Solutions - **harmonic waves** :

$$(u, v, \eta) = (u_0, v_0, \eta_0)e^{i(\omega t - \mathbf{k} \cdot \mathbf{x})}, \quad (104)$$

where ω and \mathbf{k} are frequency and wavenumber, respectively.

Algebraic system for (u_0, v_0, η_0) :

$$\begin{pmatrix} i\omega & -f & -igk_x \\ f & i\omega & -igk_y \\ -iH_0k_x & -iH_0k_y & i\omega \end{pmatrix} \begin{pmatrix} u_0 \\ v_0 \\ \eta_0 \end{pmatrix} = 0, \quad (105)$$

Dispersion equation

Condition of solvability :

$$\det \begin{pmatrix} i\omega & -f & -igk_x \\ f & i\omega & -igk_y \\ -iH_0k_x & -iH_0k_y & i\omega \end{pmatrix} = 0, \quad (106)$$

which gives :

$$\omega (\omega^2 - gH_0k^2 - f^2) = 0. \quad (107)$$

Physical meaning of solutions

3 roots of the equation correspond to

- ▶ Stationary solutions $\omega = 0$
- ▶ Propagative waves with the dispersion relation :

$$\omega^2 - gH_0 k^2 - f^2 = 0 \quad (108)$$

inertia-gravity waves.

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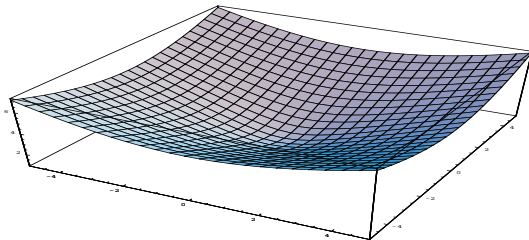
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Dispersion relation



Dispersion relation for inertia-gravity waves. $c = \sqrt{gH_0} = 1$, $f = 1$, the part with $\omega < 0$ is not presented.

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1. Demonstrate (99),
2. Obtain the **polarisation relations**, i.e. the relations between u_0, v_0, η_0 for inertia - gravity waves,
3. Calculate phase and group velocity of inertia - gravity waves,
4. Demonstrate that inertia-gravity waves **bear no** PV anomaly (PV anomaly : $q - f/H_0$),
5. Determine the spectrum of small perturbations in the 2-layer RSW model,
6. Demonstrate that PV of each layer in multi-layer RSW is

$$q_i = \frac{\zeta_i + f}{h_i}, \quad i = 1, 2, \quad \frac{d_i q_i}{dt} = 0. \quad (109)$$

Primitive equations, ocean

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$$\frac{\partial \vec{v}_h}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v}_h + f \hat{z} \wedge \vec{v}_h = - \frac{\vec{\nabla}_h \pi}{\rho} \equiv - \vec{\nabla}_h \phi, \quad (110)$$

$$\partial_t \sigma + \vec{v} \cdot \vec{\nabla} \sigma + w \rho'_s(z) = 0. \quad (111)$$

$$g \frac{\sigma}{\rho_0} = - \partial_z \phi, \quad \vec{\nabla}_h \cdot \vec{v}_h + \partial_z w = 0, \quad (112)$$

Conservation of potential vorticity - PE model

Absolute vorticity - PE model

$$\vec{\zeta}_a = \vec{\zeta} + \hat{\mathbf{z}}f, \quad \vec{\nabla} \cdot \vec{\zeta}_a = 0, \quad (113)$$

with relative vorticity :

$$\vec{\zeta} = -\partial_z v \hat{\mathbf{x}} + \partial_z u \hat{\mathbf{y}} + (\partial_x v - \partial_y u) \hat{\mathbf{z}} \quad (114)$$

Application of $\vec{\nabla} \wedge$ to PE + "hydrodynamic identity" :

$$\vec{v} \cdot \vec{\nabla} \vec{v} = \frac{1}{2} \vec{\nabla} \vec{v}^2 - \vec{v} \wedge (\vec{\nabla} \wedge \vec{v}) \quad (115)$$

→ equation for ζ_a :

$$\frac{d\vec{\zeta}_a}{dt} = \vec{\zeta}_a \cdot \vec{\nabla} \vec{v} + \frac{g}{\rho_0} \hat{\mathbf{z}} \wedge \vec{\nabla} \sigma. \quad (116)$$

Conservation of potential vorticity

$$\frac{dq}{dt} = 0, \quad q = \vec{\zeta}_a \cdot \vec{\nabla} \rho, \quad \rho = \rho_0 + \rho_s(z) + \sigma. \quad (117)$$

Démonstration.

$$\begin{aligned} \partial_t (\vec{\zeta}_a \cdot \vec{\nabla} \rho) &= (\partial_t \vec{\zeta}_a) \cdot \vec{\nabla} \rho + \vec{\zeta}_a \cdot \vec{\nabla} (\partial_t \rho) \\ &= \vec{\nabla} \rho \cdot (\vec{\nabla} \wedge (\vec{v} \wedge \vec{\zeta}_a)) - \vec{\zeta}_a \cdot \vec{\nabla} (\vec{v} \cdot \vec{\nabla} \rho) \\ &= -\vec{\nabla} \cdot (\vec{\nabla} \rho \wedge (\vec{v} \wedge \vec{\zeta}_a)) - \vec{\zeta}_a \cdot \vec{\nabla} (\vec{v} \cdot \nabla \rho) \\ &= -\vec{\nabla} \cdot (\vec{v} (\vec{\zeta}_a \cdot \vec{\nabla} \rho)) + \vec{\nabla} \cdot (\vec{\zeta}_a (\vec{v} \cdot \vec{\nabla} \rho)) \\ &\quad - \vec{\zeta}_a \cdot \vec{\nabla} (\vec{v} \cdot \nabla \rho) = -\vec{v} \cdot \vec{\nabla} (\vec{\zeta}_a \cdot \vec{\nabla} \rho). \quad (118) \end{aligned}$$



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Identities used :

$$\vec{\nabla} A \cdot (\vec{\nabla} \wedge \vec{B}) = -\vec{\nabla} \cdot (\vec{\nabla} A \wedge \vec{B}), \quad (119)$$

$$\vec{A} \wedge (\vec{B} \wedge \vec{C}) = \vec{B} (\vec{A} \cdot \vec{C}) - \vec{C} (\vec{A} \cdot \vec{B}), \quad (120)$$

and also

$$\vec{\nabla} \cdot \vec{v} = \vec{\nabla} \cdot \vec{\zeta}_a = 0$$

.

Spectrum of small perturbations - PE model

Linearised equations :

Perturbations about the state of rest : $\vec{v} = 0$ with **constant stratification** on the f - plane. Linearised equations :

$$\begin{aligned}u_t - fv + \phi_x &= 0, \\v_t + fu + \phi_y &= 0,\end{aligned}\tag{121}$$

$$\begin{aligned}\phi_z + \frac{g}{\rho_0}\sigma = 0, \quad \sigma_t + w\rho'_s &= 0, \\u_x + v_y + w_z &= 0,\end{aligned}\tag{122}$$

where u, v, w - three components of velocity perturbation, ϕ - geopotential perturbation, σ - perturbation of the profile of background density ρ_s , with $\rho'_s = \text{const.}$

Elimination of σ and w :

Elimination of σ :

$$\phi_{zt} + wN^2 = 0, \quad (123)$$

where $N^2 = -\frac{g\rho'_s}{\rho_0}$ - Brunt - Väisälä frequency

Elimination of w :

$$\begin{aligned} u_t - fv + \phi_x &= 0, \\ v_t + fu + \phi_y &= 0, \end{aligned} \quad (124)$$

$$u_x + v_y - N^{-2}\phi_{zzt} = 0, \quad (125)$$

Method of Fourier

Solutions - **harmonic waves** :

$$(u, v, \phi) = (u_0, v_0, \phi_0) e^{i(\omega t - \mathbf{k} \cdot \mathbf{x})}, \quad (126)$$

where ω et $\mathbf{k} = (k_x, k_y, k_z)$ are frequency and wavenumber, respectively.

Algebraic system for (u_0, v_0, ϕ_0) :

$$\begin{pmatrix} i\omega & -f & -ik_x \\ f & i\omega & -ik_y \\ -ik_x & -ik_y & i\frac{\omega}{N^2}k_z^2 \end{pmatrix} \begin{pmatrix} u_0 \\ v_0 \\ \eta_0 \end{pmatrix} = 0, \quad (127)$$

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Dispersion equation

Condition of solvability :

$$\det \begin{pmatrix} i\omega & -f & -ik_x \\ f & i\omega & -ik_y \\ -ik_x & -ik_y & i\frac{\omega}{N^2}k_z^2 \end{pmatrix} = 0, \quad (128)$$

which gives :

$$\omega \left(\omega^2 - \left(N^2 \frac{k_x^2 + k_y^2}{k_z^2} + f^2 \right) \right) = 0. \quad (129)$$

Physical meaning of solutions

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Three roots of this equation correspond to

- ▶ Stationary solutions $\omega = 0$
- ▶ Propagative waves with dispersion relation :

$$\omega^2 = N^2 \frac{k_x^2 + k_y^2}{k_z^2} + f^2 \quad (130)$$

Internal inertia-gravity waves : IGW.

Remark : at each fixed k_z - dispersion relation of RSW
with $c \rightarrow \frac{N}{|k_z|}$

Exercise

1. Demonstrate (116)
2. Demonstrate that PV in the PE in isopycnal coordinates :

$$q = \frac{\partial_x v - \partial_y u + f}{\partial_\theta z} \quad (131)$$

is conserved : $\frac{dq}{dt} = 0$.

3. Demonstrate the Eulerian conservation of energy in the PE, with energy density defined as :

$$e = \rho_0 \frac{u^2 + v^2}{2} + \rho g z, \quad (132)$$

where $\rho = \rho_s + \sigma$, z - (Lagrangian) position of the elementary volume of fluid.

4. Establish polarisation relations and calculate phase and group velocities of the IGW.

Euler equations for an incompressible fluid in the rotating frame without hydrostatic hypothesis :

$$\partial_t \vec{v} + \vec{v} \cdot \vec{\nabla} \vec{v} + f \hat{z} \wedge \vec{v} = -\frac{1}{\rho_0} \vec{\nabla} P \quad \vec{\nabla} \cdot \vec{v} = 0. \quad (133)$$

Linearisation ($\rho_0 = 1$) :

$$\begin{aligned} u_t - fv + P_x &= 0 \\ v_t + fu + P_y &= 0 \\ w_t + P_z = 0, \quad u_x + v_y + w_z &= 0 \end{aligned} \quad (134)$$

Solution : **inertial (gyroscopiques) waves** with dispersion relation :

$$\omega^2 = f^2 \frac{k_z^2}{k_x^2 + k_y^2 + k_z^2} \quad (135)$$

→ sub-inertial.

Full non-hydrostatic equations

$$g \frac{\rho}{\rho_0} = -\phi_z \rightarrow \frac{dw}{dt} + g \frac{\rho}{\rho_0} = -\phi_z. \quad (136)$$

Elimination of b and w :

$$b = -\phi_z - w_t, \quad -(\partial_{tt} + N^2)(u_x + v_y) + \phi_{zzt} = 0 \Rightarrow \quad (137)$$

$$u_t - fv = -\phi_x, \quad (138)$$

$$v_t + fu = -\phi_y, \quad (139)$$

$$(\partial_{tt} + N^2)(u_x + v_y) - \phi_{zzt} = 0, \quad (140)$$

Dispersion relation :

$$\omega \left[\omega^2 - \left(N^2 \frac{k^2 + l^2}{k^2 + l^2 + m^2} + f^2 \frac{m^2}{k^2 + l^2 + m^2} \right) \right] = 0 \quad (141)$$

Typically in the atmosphere and ocean

$$N^2 > f^2 \Rightarrow f^2 \leq \omega^2 \leq N^2 \quad (142)$$

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- ▶ Two dynamical entities : **waves and vortices**
- ▶ Vortices : **slow** motions related to Lagrangian conservation of PV ; zero frequency in linear approximation.
- ▶ Waves : **fast** motions
- ▶ Frequencies of wave and vortices are separated by a **spectral gap** in hydrostatic approximation.

GFD : vortices, waves, and topography

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