Wave groups and small scale variability of wave heights observed by altimeters

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7	Key Points:
8	• Wave groups contribute to small-scale fluctuations in altimeter wave heights, explaining 25 % of the variance measured by CFOSAT
10	• For the same wave height, fluctuations are larger in the presence of long and narrow- banded waves, typical of swell-dominated conditions.
12	• Altimeters smooth out scales shorter than the square root of half the H_s times the altitude, and distort spatial patterns at that scale.

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14 Abstract

Recent satellite altimeter retracking and filtering methods have considerably reduced the 15 noise level in estimates of the significant wave height (H_s) , allowing to study processes 16 with smaller spatial scales. In particular, previous studies have shown that wave-current 17 interactions may explain most of the variability of H_s at scales 20 to 100 km. As the spa-18 tial scale of the measurement is reduced, random fluctuations emerge that should be as-19 sociated to wave groups. Here we quantify the magnitude of this effect of wave groups, 20 and their contribution to the uncertainty in H_s measurements by altimeters, with a par-21 ticular focus on extreme extra-tropical storms. 22

We take advantage of the low orbit altitude of the China-France Ocean Satellite (CFOSAT), and the low noise level of the nadir beam of the SWIM instrument. Our estimate of wave group effects uses directional wave spectra measured by off-nadir beams on SWIM and signal processing theory that gives statistical properties of the wave envelope, and thus the local wave heights, from the shape of the wave spectrum.

We find that, on average, the standard deviation of H_s associated to wave groups is of the order of 5% of H_s , which is about half of the variability measured by CFOSAT over a 80 km distance. This fraction can be larger in storms and in the presence of long swells. When the estimated effect of wave groups is subtracted from the measured H_s variance, the remaining variability is strongest in regions of strong currents.

³³ Plain Language Summary

Satellite altimeters routinely provide measurements of the height of ocean waves, 34 and improved instruments or processing techniques have led to more precise and detailed 35 measurements. Here we use a combination of simulations and data from the China France 36 Ocean satellite to demonstrate that there is a limit to how accurate the measurement 37 can be due to the presence of local wave height variations, associated to the random fluc-38 tuations in the wave field known as wave "groups". There is also a limit to the horizon-39 tal size of the details that can be resolved by altimeters. Both are a function of wave height 40 and satellite altitude. 41

42 **1** Introduction

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As wind-waves impact all activities at sea, air-sea interactions and remote sensing, there is a general need for obtaining more accurate and higher resolution information about the sea state. Today, satellite radar altimetry is the most extensive source of measurements with a global coverage, providing routine estimates of the significant wave height H_s (Ardhuin et al., 2019). As these data are getting used for a wide range of applications, it is important to understand what can be measured with altimeters, at what scale and with what uncertainty.

The scale of the measurement was particularly discussed by Chelton et al. (1989), who introduced the concept of *oceanographic footprint* which contains the sea surface points that contribute to the measurement of sea level and wave height. This footprint is a disc of radius

$$r_C = \sqrt{\frac{2hH_s + 2\delta_R}{1 + h/R_E}} \tag{1}$$

where *h* is the satellite altitude, R_E is the Earth radius, H_s is the wave height and the range resolution $\delta_R = c/(2B)$ is defined by the radar bandwidth *B* and the speed of light *c*. We note that all Ku-band altimeters have used B = 320 MHz giving $\delta_R = 0.47$ m, and B = 500 MHz on SARAL-AltiKa gives $\delta_R = 0.32$ m, so that r_C is always larger

 $_{59}$ than 1 km. Because low Earth orbit satellites fly over the ocean at a speed around 7 km/s,

averaging altimeter data over 0.05 s corresponds to a spatial average over 350 m and thus
 does not change much the effective footprint of the measurement.

The apparent noise in such data is generally attributed to the uncertainty of the 62 measurement, and has led users of altimetry data to take longer averages of H_s , over 1 63 to 10 s, corresponding to a distance that spans 7 to 70 km. While it effectively reduces 64 noise, such averaging also blurs potentially interesting features, in particular the peaks 65 of storms, coastal gradients (Passaro et al., 2021), and the signature of surface currents 66 (Quilfen & Chapron, 2019). Away from surface current gradients and coastlines, sea states 67 are uniform over scales of the order of 70 km (Tournadre, 1993). Still, within these scales, 68 the random nature of the wave field is another source of expected geophysical variabil-69 ity. Theoretical analysis, in situ time series and airborne remote sensing show that small 70 scale variations in H_s contain a signature of wave groups that can be estimated from the 71 wave spectrum (Arhan & Ezraty, 1978; Tayfun & Lo, 1989), and is the result of the lin-72 ear superposition of many wave trains. Wave groups have typical time scales of a few 73 tens of seconds to a few minutes, that translate to spatial patterns at scales of a few kilo-74 meters, hence around the possible resolution limit of altimeters, of the order of r_C . At 75 larger scales, non-linear wave-wave interactions should contribute to fluctuations at scales 76 10 to 20 minutes, with spatial scales around 10 km, that should be important for wave 77 growth (Lavrenov, 2001) and may contain information on the wave period (Badulin, 2014). 78 Co-location of altimeter, buoy and model data with wave heights from 1 to 8 m, has been 79 used to estimate a typical uncertainty of 1 s averaged altimeter data around 7% for $H_s >$ 80 2 m (Dodet et al., 2022). Understanding what makes up this uncertainty will help ex-81 trapolate uncertainties to higher values of H_s , providing a better understanding of the 82 climatology of sea state extremes. 83

In the present paper we focus our analysis on the fluctuations of H_s associated to 84 wave groups and its contribution to Delay-only altimeters that provide the existing ref-85 erence time series for wave climate analysis (Young et al., 2011; Dodet et al., 2020). The 86 main question that we wish to answer is: how much wave groups contribute to the vari-87 ability in H_s measurements? For this we take advantage of the unique opportunity pro-88 vided by the SWIM instrument onboard the China-France Ocean Satellite (CFOSAT). 89 SWIM provides both directional wave spectra from which we compute the spectrum of 90 the wave envelope that contains wave groups, and along-track nadir altimetry. Our anal-91 ysis uses SWIM data over the globe for two full years 2020 and 2021. 92

We start with two illustrative and contrasting examples in section 2, before pro-93 viding results for the globe in section 3. Discussions and conclusion follow in section 4. 94 A side question that we had to address is: how does an altimeter measure H_s over a re-95 alistic surface that contains local perturbations associated with wave groups? For this 96 we used a very simplified simulated altimeter with numerical results shown in section 97 2 and an analytical derivation in Appendix A. Those results suggest that altimeters re-98 port a particular kind of average H_s over an radius that is close to $r_C/2$ while ampli-99 fying perturbations around $r_C/4$ and ignoring perturbations located right at the nadir. 100

¹⁰¹ 2 One particular storm and a theory of wave groups

As described in Hauser et al. (2017), the instrument SWIM is a Ku-band wave scat-102 terometer that illuminates successively 6 incidence angles $(0^{\circ}, 2^{\circ}, 4^{\circ}, 6^{\circ}, 8^{\circ})$ and 10° . The 103 nadir beam (0°) works as all previous Poseidon radar altimeters and provides an esti-104 mate of H_s every 0.22 s, using an average over 0.055 s. As a result, the nadir beam data 105 is expected to be similar to data from previous Ku-band altimeters, such as Poseidon-106 3B on Jason-3, with the specific difference given by a lower data rate (5 Hz instead of 107 20 Hz for the native estimates of H_s) and a different measurement geometry associated 108 to a rather low orbit. In principle, the low orbit altitude h = 519 km of CFOSAT make 109 it possible to resolve smaller scale variations of H_s as r_C is reduced by a factor 1.4 com-110

pared to the Jason satellites that orbit at 1340 km altitude. The low noise level of the satellite and specific processing of the SWIM instrument also contribute to its capability to resolve smaller scales (Tourain et al., 2021).

The off-nadir beams use the concept of the wave spectrometer (Jackson et al., 1985) 114 based on a real-aperture radar and the normalized radar cross-section (NRCS) sensitiv-115 ity to local surface slope at near-nadir incidences, providing estimates of the directional 116 wave spectra (with a 180° ambiguity in direction). The CNES mission center (or CFOSAT 117 Wind and Waves Intrument Center - CWWIC) provides Level 2 products both for the 118 nadir beam and the off-nadir beams 6° to 10°. The off-nadir L2-CWWIC products con-119 sist of 2D wave spectra provided in 12 directions from 0 to 180° and in 32 wavenumbers 120 and constructed from overlapping of antenna scans over 180° (on each side of the track) 121 over boxes of about 70 km by 90 km. In order to allow comparison, the nadir product 122 is resampled by averaging values of Hs over the box size, its variation at this scale is quan-123 tified by taking its standard deviation over the same distance, $std(H_s)$. We will partic-124 ularly investigate this quantity in the present paper. 125

Alternatively, the Ifremer Waves and Wind Operational Center (IWWOC) is in charge 126 of implementing a different processing and provides its own Level 2 products for off-nadir 127 beams. These products are referred to as L2S products and consist of 1D wave modu-128 lation spectra, one for each antenna scans. Whereas the L2-CWWIC product uses the 129 nadir H_s to rescale the spectrum, the L2S product is based on a theoretical modulation 130 transfer function to transform the NRCS spectra into surface elevation spectra (Jackson 131 et al., 1985). Also the L2-CWWIC product uses a very conservative maximum wavelength 132 of 500 m in order to avoid amplifying noise, where the L2S product does not use such 133 a fixed value for the maximum wavelength. As a result, L2S spectra may capture the very 134 long waves produced in the most severe storms. 135

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2.1 Wave height variability in Storm Dennis

On February 14^{th} 2020, the European windstorm Dennis, which became one of the 137 most intense extratropical cyclones ever recorded, underwent through its explosive in-138 tensification in the middle of the North Atlantic. Around 9:10 UTC that same day, it 139 happened to cross track with CFOSAT, whose altimeter recorded H_s values up to 19.7 m 140 for the native sampling, and 17.9 m for the 1 Hz sampling (averaging over 1Hz). Fig. 1.a 141 shows a model snapshot of H_s in the north Atlantic and the corresponding descending 142 track of CFOSAT, while Fig. 1 shows the altimeter wave height H_s values for the three 143 different samplings : native (4.5 Hz), 1 Hz and box averaged. 144

On the periphery of the storm, where the average H_s is around 10 m, we were struck 145 by the factor two difference in std(Hs). Because this difference is not localized but spans 146 more than 420 km (1 minute of data), we hypothesize that the spatial variability in winds 147 and currents forcing, that are known to cause variability in H_s (Abdalla & Cavaleri, 2002; 148 Ardhuin et al., 2017) may not play a dominant role in this difference. In particular, re-149 gions of high current variability are usually much more localized (Quilfen & Chapron, 150 2019). Our working hypothesis is that this variability of H_s may be dominated by fluc-151 tuations in wave heights associated to wave groups. These fluctuations have different mag-152 nitudes and spatial scales which can be estimated from the directional wave spectrum. 153 Hence CFOSAT is a unique instrument for studying this effect as we have both H_s vari-154 ability along the satellite track and directional wave spectra. In the following, we will 155 illustrate the expected signature of wave groups for the two sea state conditions that cor-156 respond to the particular SWIM boxes highlighted in cyan and magenta. We note that 157 in these two examples, the H_s values obtained from the sum of the L2S spectrum are 158 around 7.5 m, which is lower than the 9 m given by the nadir beam and used in the L2 159 product to rescale the specturm energy. 160



Figure 1. Map of wave heights in the North Atlantic at 09:00 on 14 February 2020, as provided by the model hindcast of (Alday et al., 2021), overlaid with circles located at the center of SWIM box estimates for the L2-CWWIC wave spectra. Circles are sized by the L2-CWWIC H_s estimate and color corresponds to std(Hs).

2.2 Variability of H_s and envelope spectrum

The patterns of individual waves vary with the shape of the wave spectrum, as il-162 lustrated in Fig. 2. A key difference between the north-side (left column) and south-side 163 (right column) of storm Dennis is that the south-side has a longer peak wavelength around 600 m, and a more narrow spectrum, in particular in directions. The smaller width in 165 directions gives longer wave crests while the smaller width in wavenumber magnitude 166 gives a larger number of waves in the succession of large waves, known as groups. Us-167 ing the envelope of the sea surface we can quantify the size of these groups (Arhan & 168 Ezraty, 1978; Longuet-Higgins, 1984; Masson & Chandler, 1993), and their contribution 169 to the spatial variability of H_s at all scales. 170

Let ζ_a be the complex surface such that $\zeta = \operatorname{Re}(\zeta_a)$ is the free surface. The en-171 velope η of the signal is defined by $\eta = |\zeta_a|$. This defines a local amplitude of the sig-172 nal, that does not contain the small scale crest-to-trough (positive to negative) varia-173 tions of the original surface. Variations of the envelope contain scales much larger than 174 the wavelengths, including scales that are comparable to, or larger than the footprint of 175 a satellite altimeter. Hence wave groups may contribute to the fluctuations of H_s recorded 176 by the nadir beam of SWIM, as indicated on Fig. 1.b. We will now attempt to quantify 177 that contribution. 178

¹⁷⁹ We thus define a local H_s as 4 times the standard deviation of the sea surface over ¹⁸⁰ a distance r_a . Using the Gaussian approximation of the distribution of sea surface el-¹⁸¹ evations leads to an envelope following a Rayleigh distribution of parameter $H_s/4$, thus ¹⁸² the mean of the envelope is $H_s\sqrt{\pi}/(4\sqrt{2})$. By extension, $4\sqrt{2/\pi}$ times an average of the ¹⁸³ envelope over the spatial scale r_a is equal to our local H_s .

Therefore we can write,

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$$H_{s,r_a}(x,y) = 4\sqrt{\frac{2}{\pi}}(\eta \otimes g_{r_a})(x,y)$$
⁽²⁾

where \otimes is the convolution operator and g_{r_a} is a filtering kernel of radius r_a .



Figure 2. Left column corresponds to our chosen northern CFOSAT box, and right column to the southern box. Top line: L2S wave spectra $E(k, \theta)$ as provided by IWWOC with k the wavenumber. Middle line: simulated surface elevation maps generated from the wave spectra using random phases, bottom line: envelope of the surface elevation.

In order to understand how much wave groups may contribute to H_s fluctuations in satellite data, we have to address two questions: First, what are the scales affected by wave groups? and second, what are the scales of the H_s variation that are resolved by satellite altimeters?

One simple way to quantify the different scales present in the envelope is to compute its spectrum. The most simple theoretical result comes directly from the theory of Fourier transforms that gives the spectrum of a product of functions as the convolution of the Fourier transforms. In our case, the envelope squared is the product of the elevation by its complex conjugate, and this is true for spectra in one or two dimensions. For waves in one dimension with wavenumber k, the spectrum of the envelope squared $\Psi_2(k)$ is the convolution of the spectrum of the one-sided surface elevation spectrum E(k)by itself,

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$$\Psi_2(k) = 8 \int_0^\infty E(u)E(u+k)\mathrm{d}u,\tag{3}$$

and we note that $\Psi_2(k)$ is also single-sided.

In practice people have rather studied the variations of H_s and not that of H_s^2 . Although the details of the theory are more complex, the important result is that, for low frequencies, the spectrum of the envelope $\Psi(k)$ has the same shape as the spectrum of the envelope squared $\Psi_2(k)$ (Rice, 1944). More specifically, Tayfun and Lo (1989) have showed that a good approximation for the spectrum of the envelope is given by

$$\Psi(k) = \frac{8 - 2\pi}{H_s^2} \Psi_2(k) \tag{4}$$

This same result is valid for spectra in two dimensions. We now consider the doublesided wave spectrum $E(k_x, k_y)$, defined for (k_x, k_y) in the entire wavenumber plane and centrally symmetric, the region of the envelope spectrum for $k \ll k_p$ is identical to

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$$\Psi_2(k_x, k_y) = 8 \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} E(u, v) E(u + k_x, v + k_y) \mathrm{d}u \mathrm{d}v,$$
(5)

in which Ψ_2 is also double-sided.

From eq. (2), the spectrum of $H_s^{r_a}$ is

$$\Psi_{H_s,r_a}(k_x,k_y) = 4\sqrt{\frac{2}{\pi}} \cdot \frac{8-2\pi}{H_s^2} \cdot \Psi_2(k_x,k_y)G_{r_a}(k_x,k_y)$$

with G_{r_a} the square of the Fourier transform of the altimeter filtering kernel g_{r_a} . We can thus estimate the standard deviation of H_{s,r_a} in altimeter measurements along the satellite track for segment of length L_1 , by integrating the expected variance for $k_x > k_1$, with $k_1 = 2\pi/L_1$ and the x-axis taken in the along-track direction. The group-induced variation of H_s is thus equal to

$$\operatorname{var}(H_{s,r_a}, L_1)_{wg} = 16\sqrt{\frac{2}{\pi}} \frac{4-\pi}{H_s^2} \int_{-\infty}^{\infty} \int_{k_1}^{\infty} \Psi_2(k_x, k_y) \cdot G_{r_a}(k_x, k_y) \mathrm{d}k_x \mathrm{d}k_y.$$
(6)

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We now need to estimate r_a and G_{r_a} for the comparison with altimeter data.

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2.3 Altimeter measurements over varying wave heights

Satellite altimeters transmit radar pulses that are reflected from the sea surface, 222 and they measure the backscattered power as a function of time. This variation of re-223 ceived power as a function of time is known as the waveform, and it is typically aver-224 aged over a few hundred pulses spanning about 0.05 s in order to reduce the speckle noise 225 (Quartly et al., 2001). Time can be transformed to distance from the satellite, or range 226 R, and the waveform effectively contains information on the statistical distribution of 227 ranges around the mean satellite altitude h. As shown by Brown (1977), assuming a uni-228 form ocean reflectivity and broad radar antenna pattern, the waveform is a area-weighted 229 histogram of the ranges. Over a flat sea surface, this histogram is a Heaviside function 230 because the part of the ocean surface with ranges between R and $R + \delta_R$ is an annu-231 lus of radius $r = \sqrt{R^2 - h^2}$ centered on the nadir point, with an area $2\pi R \delta_R$ that is 232 almost constant as long as $R \approx h$. In the presence of waves, a negative surface ele-233 vation $z = \zeta$ will shift the range to a higher value, sharing the range position of sea sur-234 face elements located at z = 0 and further away from nadir. Given the very small in-235 cidence angles, the change in range is $\Delta_R = -\zeta$. For a Gaussian distribution of ζ with 236

standard deviation $\sigma_H = H_s/4$, the presence of waves gives a smoothing of the histogram, hence the theoretical waveform (Brown, 1977),

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$$w_B(R,\sigma_H) = \frac{1}{2} \left[1 + \operatorname{erf}\left(\frac{R-h}{\sqrt{2}\sigma_H}\right) \right].$$
(7)

When "retracking" altimeter data, eq. (7) is inverted, and H_s is estimated to be 240 4 times the σ_H of the theoretical waveform that best fits the data. In practice the the-241 oretical waveform also includes effects of the antenna patterns, possible mispointing, and 242 different fitting methods have been developed to reduce the effect of noise or spurious 243 echoes in the measured waveform (Passaro et al., 2014; Tourain et al., 2021). Another 244 important assumption needed to obtain the Brown waveform is that the sea state is ho-245 mogeneous within the footprint. We thus have to discuss what sets the scale of the foot-246 print, or more precisely where are the points on the sea surface that give the distinctive 247 shape of the waveform and allow the fit to distinguish different wave heights. 248

Compared to a flat sea surface, the elevation ζ at a distance r from the nadir point will change the range R of the surface point and make it look as if it was located at a different distance $r + \delta_r$, so that points from different locations on the sea surface will map to the same range R. This is the same "range bunching" or "overlay" or "surfboard effect" that is common to all radar systems (Peral et al., 2015). Following Chelton et al. (1989) we can estimate the apparent horizontal displacement. For a satellite altitude hand using $\zeta \ll R$, the calculation for a flat mean sea surface gives

$$\delta_r \simeq \sqrt{r^2 - 2h\zeta} - r. \tag{8}$$

For a spherical Earth of radius R_E , ζ should be replaced by $\zeta/(1 + h/R_E)$.

In the particular case where $\zeta = -H_s$ and r = 0, δ_r is the radius r_C of oceanographic footprint as defined by Chelton et al. (1989), and given in eq. (1), when the range resolution δ_R is neglected compared to H_s . For $H_s = 10$ m, and h = 519 km, this gives $r_C = 3.3$ km.

We may give the following interpretation of r_C . At points located at r_C from nadir 262 (i. e. at the edge of the "Chelton footprint") there is a 99.997% probability that $\zeta <$ 263 H_s and that these points contribute to the waveform at ranges R > h, i.e. in the sec-264 ond half of the rising part of the waveform. This definition means that, in practice, points 265 outside of the footprint do not contribute to the middle part of the waveform that most 266 contributes to the fit of σ_H and the estimation of H_s . The definition of the footprint size 267 given by Chelton et al. (1989) is thus truly a maximum size. In practice, at a distance 268 $r_C/2$ from nadir, the probability to contribute only to the second half of the waveform 269 is already 85%, so that the radius of the footprint that contains half of the points that 270 contribute to the middle region of the waveform is of the order of $r_C/2$. 271

If wave heights vary as a function of distance to nadir, then the waveform does not follow exactly the Brown form, as detailed in Appendix A. As different regions of the waveform contain different regions of the sea surface, one could imagine fitting different parts of the waveform to measure variations in H_s as a function of distance from nadir. The theoretical limit to this capability is the blurring due to range bunching over a distance of the order $r_C/2$. Speckle noise is another limiting factor, and probably the main one in practice.

Based on the analysis in Appendix A we expect that variations of H_s at scales much smaller than $r_C/4$ will be smoothed out in altimeter data, whereas variations at scales much larger than $r_C/2$ have no effect on the waveform that will follow the Brown shape. For our analysis of CFOSAT data we will define an "effective altimeter radius" r_a such that the variance of H_s associated to the random fluctuations of the envelope filtered with a Gaussian filter of standard deviation r_a , is the same as that produced by an altimeter. The actual shape of the "altimeter filter" is discussed in Appendix A.

2.4 Estimation of the equivalent r_a scale for an idealized altimeter

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We have simulated the sampling of our simulated sea surface by a highly simpli-287 fied altimeter. We thus neglect radar noise, speckle and variations in ocean backscatter, 288 and compute simulated waveforms as histograms of the number of discrete pixels as a 289 function of range R discretized with the same resolution $\delta_R = 0.47$ m used in actual 290 SWIM data. The histogram is computed for a finite region of size r_C by r_C centered at 291 the nadir point. The value of H_s for each simulated histogram is taken to be the H_s of 292 the theoretical waveform given by eq. (7) that best fits the simulation for R varying from 293 $h-H_s$ to $h+H_s$, using an L2-distance. As detailed in Appendix A, the altimeter makes 294 a much more complex measurement than a simple Gaussian smoothing of the H_s field. 295 We also expect that more realistic waveforms and different fitting procedure may yield 296 some differences. 297

Taking the simulated sea surface from Fig. 2, we compare a map of simulated altimeter data in Fig. 3 with maps of local wave heights, smoothed on different scales.



Figure 3. Maps of wave heights obtained by either simulating altimeter processing (top left) or smoothing the envelope with a Gaussian filter of standard deviation r_0 varied from $r_C/5$ to $r_C/2$. In this example, $r_C = 2785 m$ thus r_0 values are respectively 557, 619, 696, 928, and 1392 m. In practice the smoothing is applied in a finite box of size $4r_C$ by $4r_C$.

As expected, the large scales of the envelope, those that persist in the bottom-right panel of Fig. (3), match the large scales of the simulated altimeter data. From a quantitative point of view, the standard deviation of the simulated altimeter data, here 0.63 m, is of the same order as the standard deviation of actual SWIM measurements over the same SWIM box. We also note that this value is very close to that obtained for a filtering of the envelope at a scale $r_a = r_C/4.5$.

Looking at the top left panel of Fig. 3 it is clear that the map of retracked H_s contains much smaller features than the envelope smoothed with $r_a = r_C/4.5$. All of these are spurious amplification of envelope perturbations that happen to be at the right distance from nadir, around $r_C/2$, as explained in Appendix A. As a result, maxima of H_s given by the altimeter are not located at the true H_s maxima but slightly displaced by a distance of the order of $r_C/2$. A striking example in Fig. 3 is the region of waves higher than 11 m around x = 33 km, y = 8 km. The altimeter gives a local minimum where the true wave height is maximum, and a round halo of maxima surrounding that point. Conversely a ring-shaped maximum in the envelope, such as around x = 19 km, y =9 km, gives a strong maximum in the simulated altimeter data. We expect that sensitivity may vary with the actual retracking procedure, and this discussion is beyond the scope of the present paper.

Even though the patterns do not exactly coincide, we will now assume that the sampling of the sea surface by the altimeter is equivalent, in terms of variability of H_s to filtering the envelope with a Gaussian of standard deviation $r_a = r_C/4.5$.

2.5 Generic sea states, envelopes and H_s variability

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Based on our analysis, we expect that SWIM measurements of H_s are contaminated 322 by wave group structures at scales of the order of a few kilometers, following the vari-323 ation of r_C with wave height and satellite altitude. As illustrated by the two examples 324 with different spectral width, we note that for the same wave height, a wider spectrum 325 leads to smaller scales of wave groups, part of which scales are smoothed away by the 326 altimeter footprint and therefore not resolved. For a narrower spectrum, wave groups 327 have larger scales and amplitudes and a larger contribution to the variability of wave heights 328 estimated by an altimeter. 329

For a simple quantitative analysis we may consider the more simple case of waves propagating in only one direction, with a sea surface ζ distributed with the normal law $\mathcal{N}(0, \sigma_H = H_s/4)$ with a one-sided Gaussian spectrum (defined for k > 0)

$$E(k) = \frac{H_s^2}{16\sigma_k \sqrt{2\pi}} e^{-(k-k_p)^2/(2\sigma_k^2)}.$$
(9)

The spectrum of the envelope is also Gaussian and the H_s PSD writes (for a one-sided spectrum),

$$\Psi_{H_s}(k) = \frac{4 - \pi}{\pi \sqrt{\pi} \sigma_k} \cdot H_s^2 e^{-k^2/(4\sigma_k^2)}.$$
 (10)

Wave groups contain wavelengths larger than π/σ_k , with a constant spectral density near k = 0. Around k = 0, the value of the Hs spectrum is $0.15H_s^2/\sigma_k \text{ m}^2/(\text{rad/m})$.

Figure 4 presents one dimensional wave spectra - in solid lines - of two typical sea states with same $H_s = 3.11$ m and their associated H_s spectra $\Psi_{H_s}(k)$ - in dashed lines. The blue spectrum is a JONSWAP spectrum (Hasselmann et al., 1973) with a peak period of 8 s and a peak enhancement factor $\gamma = 3.3$ that represents a moderate windsea. The red spectrum is a narrow Gaussian spectrum with a peak period of 14 s and $\sigma_k = 0.002$ rad/m, typical of swell conditions in the open ocean.

The altimeter smoothing function $G_{r_a} = \exp(-k^2 r_a^2)$ allows to define a cut-off wave number $k_a = \sqrt{\pi}/(2r_a)$. As shown in Fig. 4 the wavelengths in altimeter-filtered envelopes, larger than the associated wavelength cut-off $2\pi/k_a$ (in black dash-dotted line), are large compared to the shortest wavelengths contained in the wave groups (of order $\pi\sqrt{2}/\sigma_k$ and represented by the red and blue vertical dash-dotted lines).

Applying the one dimension version of eq. (6) gives the variance of altimeter-estimated H_s as the shaded areas in Fig. 4. For a Gaussian spectrum, in cases where the altimeter filter scale is large enough not to be concerned about the shortest scales, this area is approximately k_a times $\Psi_{H_s}(k=0)$ the H_s PSD level at k=0. This gives a standard deviation of H_s of the order of $0.39\sqrt{k_a/\sigma_k}H_s$, which is 0.40 H_s for the one-dimensional swell example of Fig. 4.

For a generic one-dimensional wave spectrum E(k), the width σ_k should be replaced by the bandwidth $\Lambda_k/(2\sqrt{\pi})$, with Λ_k defined like the usual frequency bandwidth



Figure 4. Example of two wave spectra - solid lines - in one dimension and the corresponding spectra of H_s - dashed lines -, for typical swell conditions in the open ocean in red, and typical moderate wind windsea in blue. Because the fluctuations of H_s are filtered by the altimeter with the function $G_{r_a}(k)$ - dashed magenta and cyan lines -, the actual measured variance of H_s is the shaded area, in purple for the windsea and pink for the swell. The vertical black line is the equivalent altimeter cut-off wavenumber at $k = k_a$, whereas the vertical red and blue lines represents the width of the H_s spectra.

 Λ (Saulnier et al., 2011),

$$\Lambda_k = \frac{H_s^4}{256 \int_0^\infty E^2(k) \mathrm{d}k}.$$
(11)

For a JONSWAP spectrum, $\Lambda_k = 1.3k_p$ and the standard deviation of H_s for the wind sea case above is $0.1H_s$.

For waves in two dimensions, the directional spread of the energy leads to a further reduction of the variability of H_s . Using the spectrum of the envelope at k = 0we may define a two-dimensional spectral bandwidth Λ_2

$$\Lambda_2^2 = \frac{H_s^4}{256\Psi_2(k_x = 0, k_y = 0)/8} = \frac{H_s^4}{256\int_{-\infty}^{\infty}\int_{-\infty}^{\infty}E^2(k_x, k_y)\mathrm{d}k_x\mathrm{d}k_y}.$$
 (12)

When r_a is large enough we can consider that Ψ is constant for $k < 1/r_a$, which gives

$$std(H_s)_{wg} \simeq \frac{\sqrt{4-\pi}}{4\Lambda_2} H_s \sqrt{\pi/r_a^2 - 2k_1/r_a}.$$
 (13)

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2.6 Summary of methodology

³⁶⁹ **3** Results at the global scale

Beyond the particular case of storm Dennis, with very large wavelength and narrow spectra that may lead to a dominant effect of wave groups in H_s variability, how important are wave groups in general, and how important can they be compared to other known sources of H_s variability, including winds and currents?

We have thus applied eq. (6) to the full SWIM L2S archive for the years 2020 and 2021, and estimated for each directional wave spectrum defined in a SWIM L2 box the expected value of $\operatorname{std}(H_s)_{wg}$ due to wave groups: we compute the 2D envelope spectrum from the convolution of the wave spectrum and integrate the expected variance of H_s from a minimum wavenumber k_1 in the along-track direction. We verified that using only



Figure 5. Left column corresponds to our chosen northern CFOSAT box, and right column to the southern box. Top line: envelope squared spectrum $\Psi_2(k_x, k_y)$ from convolution. Middle line: spectrum of $H_s^{r_a} \Psi_{H_s, r_a}$ (including the equivalent altimeter filtering). Bottom line: 1D along-track spectrum obtained by integrating over the cross-track axis, in cyan for the northern box and magenta for the southern box.

the envelope value at $k_x = 0, k_y = 0$, which is equivalent to using eq. (13) gives similar results.

This gives more than 2.4 millions of $\operatorname{std}(Hs)_{\operatorname{wg}}$ values. Because $\operatorname{std}(H_s)$ is generally proportional to the average of H_s in a box, we have shown in Fig. 7.a the distribu-



Figure 6. Values of H_s , averaged over boxes, and std(Hs); as provided in the L2-CWWIC as a function of sampling time (UTC), for the CFOSAT track shown in Fig. 1. The first measurements are located at 68° N, and the last are at 35°N. The box highlighted in cyan is at 62°N, and the one in magenta is at 44°N.

tion of $\operatorname{std}(H_s)/\operatorname{mean}(H_s)$ and, for comparison, $\operatorname{std}(H_s)_{wg}/\operatorname{mean}(H_s)$ in Fig. 7.b. Note that the colorbar have different scales so that our estimate of the contribution of wave groups is, on average, half of the measured $\operatorname{std}(H_s)$. Both figures have some common patterns with a general increase from the west to the east of the ocean basins consistent with a dominance of swells in the east with longer wavelengths and narrower spectra. Obviously, having divided by $\operatorname{mean}(H_s)$ also tends to increase the values where H_s is small in enclosed seas and near the Equator, even when removing values for $H_s < 1.5$ m.

We may also remove the contribution of wave groups to look at the other sources 390 of variability in H_s . As shown in Fig. 8, the standard deviation of H_s corrected for the 391 effects of wave groups contains a background level of 0.1 to 0.2 m, probably associated 392 to speckle noise, and stronger localized values up to 0.3 m. These large values are co-393 located with regions of strong ocean circulation mesoscale variability. These same regions 394 match the location of strong H_s gradient when the along-track data as been de-noised 395 using an Empirical Mode Decomposition (EMD) as previously demonstrated by Quilfen 396 and Chapron (2019). We can use this correspondence to interpret the EMD denosing pro-397 cedure as a removal of both speckle noise, which has nothing to do with the variability 398 of H_s , and wave group effects that are related to a true variability of H_s . 399

400 4 Discussion

The accurate estimation of wave group contributions critically depends on the ac-401 curacy of the spectral shape, in particular the directional width and wavenumber width. 402 Because of the hard wavelength cut-off in the L2 product we had chosen to work with 403 the L2S spectra. Redoing the global analysis with the L2 product generally reduces the 404 expected effect of wave groups, as shown in Fig. 9. We note that a validation of spec-405 tral width from the L2 product was performed by Le Merle et al. (2021), who found that 406 SWIM L2 generally overestimate spectral width compared to buoy data. No such anal-407 ysis has been performed for the L2S product. It would be also interesting to know how 408 accurate could be the estimation of $\operatorname{std}_{wq}(H_s)$ estimated from model spectra, for the ap-409 plication to other satellite mission that do not measure the wave directional spectrum. 410 Such a study should be careful about the spatial resolution of SWIM L2 spectra that are 411 averaged over 80 km. This averaging is expected to produce a broader spectrum that 412



Figure 7. Map of the average of $\operatorname{std}(H_s)/\operatorname{mean}(H_s)$ - upper panel - and $\operatorname{std}_{wg}(H_s)/\operatorname{mean}(H_s)$ -lower panel - for the years 2020 and 2021 for all the SWIM L2-IWWOC boxes with a H_s above 1.5 m. With the wave group contribution $\operatorname{std}_{wg}(H_s)$ estimated from SWIM L2S spectra

could underestimate the effect of wave groups. Indeed, during the concatenation over
a 80 km box, energy in two neighboring wave azimuths may not be present at the same
location but observed in different parts of the swath, some kilometers apart. This is particularly true for the higher incidence beams where that distance might be up to 50 km.

Paragraph about other missions (discussing impact of orbit altitude) ... We note
that the lower orbit of CFOSAT compared to Jason-3, for example, makes it easier to
resolve wave groups in the SWIM data. Maybe add figure with map of Hs "as seen by
Jason" compared to CFOSAT.

⁴²¹ Delay-only altimeters studied here have a noise in their H_s estimate that is prob-⁴²² ably dominated by the speckle noise in the waveforms (Sandwell & Smith, 2005; Quar-⁴²³ tly et al., 2019). Doppler processing of recent altimeter instruments starting with Cryosat-⁴²⁴ 2 and Sentinel-3 can strongly reduce this speckle noise by forming and combining inde-⁴²⁵ pendent looks of the same sea surface (Egido et al., 2021). It will therefore be interest-⁴²⁶ ing to study the effect of wave groups in these measurements of wave height and sea level. ⁴²⁷ If the Doppler induced by orbital velocities is negected, the delay-Doppler measurement



Figure 8. Residual standard deviation of H_s , in meters, after removing the effect expected from wave groups based on L2S SWIM spectra.

is similarly based on the convolution of a surface elevation distribution with a flat sur-428 face response (Ray et al., 2015). Only the flat surface response is different from the delay-429 only processing, and the details of the effect will depend on the details of how the dif-430 ferent Doppler beams are stacked. We generally expect that the blurring effect will now 431 be confined to the direction perpendicular to the track, with maximum effect of a H_s 432 perturbation located off the satellite track, possibly also at a distance of the order of $r_C/2$. 433 Because Delay-Doppler altimeters can actually resolve the along-track variability caused 434 by wave groups instead of averaging it, we expect that H_s fluctuations caused by wave 435 groups are larger in Delay-Doppler altimetry, together with their spurious effect on sea 436 level estimates. This may explain the relative smaller reduction of $std(H_s)$ at large H_s 437 which is found when Doppler resolution is enhanced to reduce the speckle effect, and the 438 typical values of $std(H_s)$ for Delay-Doppler Sentinel 3A data which is around 0.7 m for 439 $H_s = 7$ m (Egido et al., 2021), twice the typical value for SWIM data. This will be the 440 topic of further studies. 441

Up to now, the uncertainty of satellite measurements has been determined by the 442 triple-collocation method (Abdalla et al., 2011; Dodet et al., 2022), with the practical 443 result that the uncertainty of 1 Hz (TBC) altimeter data is of the order of 7% of H_s . How-444 ever, that error contains representation errors (the co-located in situ data does not sam-445 ple the same space and time frame), and cannot be extrapolated beyond the range of 446 the co-located dataset, typically wave heights below 8 m. So what can we say about the 447 largest measured wave heights of 20.1 m (Hanafin et al., 2012)? Can we use the mea-448 sured variability of H_s , for example the 5 Hz or 20 Hz data that is used to make a 1 Hz 449 average, to refine our estimate of the uncertainty of this average? In the present paper 450 we have shown that wave groups are responsible for random fluctuations in H_s estimates, 451 that are generally proportional to H_s but with an effect that depends on the bandwidth 452 of the spectrum, which is generally narrower for larger wave periods. As a result the vari-453 ability associated to wave groups can be the dominant source of fluctuations in H_s mea-454 surements for severe storm conditions. Even though the measurement fluctuations are 455 weakly correlated to the actual wave height variations (as demonstrated in Fig. 3) their 456



 $std(H_s)/H_s$ estimated L2

Figure 9. Same as Fig. 7.b, with $\operatorname{std}_{wg}(H_s)$ estimated from CWWIC-L2 wave spectra instead of IWWOC-L2S spectra.

magnitudes are strongly correlated. Hence the measured fluctuation $std(H_s)$ contains 457 both uncorrelated speckle noise effect, that can be expected to be reduced by $1/\sqrt{N}$ when 458 averaged from N Hz to 1 Hz, and a true geophysical spatial variability associated to wave 459 groups (and variable fetch, currents, etc.) that will partially average out. We expect that 460 an uncertainty model for averages of H_s measurements may take into account wave groups 461 explicitly. In the case of SWIM, directional wave spectra can be used to separate the ac-462 tual variability of the 5 Hz data into wave group effects and noise plus other geophys-463 ical effects. For other altimeters, one may use empirical correlations between spectral bandwidth, wave height and wind speed. For this information to be useful for a theoretically-465 based uncertainty estimate, which is much needed for wave heights above 8 m, one may 466 extend the parameterization of speckle effects proposed in Appendix A.3, to the actual 467 target waveform and cost function used in the retracking algorithm. 468

The present work should be useful for the exploration of the resolution limits of 469 satellite altimeters and other remote sensing systems that use radar or optical imagery 470 (Kudryavtsev et al., 2017). As processing methods are refined to produce higher reso-471 lution near the coast (Passaro et al., 2021) and the ice edge (Collard, 2022), some of the 472 high resolution data will be dominated by wave groups. The associated variance may pro-473 vide some constraint on the shape of the directional wave spectrum, but its determin-474 istic values are probably of little value for most applications as groups will travel at speeds 475 of the order of 10 m/s and persist for only a few minutes. The contribution of wave groups 476 to the variability of wave heights measured by altimeters is thus a real effect that con-477 tains part of the true variability at the scale of the altimeter footprint. Methods devel-478 oped to remove noise in the data, such as the data-driven Empirical Mode Decomposi-479 tion (EMD) used by (Quilfen & Chapron, 2019) actually remove the effect of wave groups. 480 Such data should thus be used with caution when studying the variability of wave heights 481 at the smallest scales. 482

In locations where H_s varies sharply such as over coral reefs, mud banks or across the sea ice edge, the high resolution wave heights will contain other effects, and these are particularly interesting. Some caution should be used when interpreting these sharp gradients. As we have found out, the maximum wave height will generally be displaced from the location of its true maximum. This displacement is smallest for the SWIM instrument, thanks to the low orbit of CFOSAT, which makes it a particularly interesting instrument for studying small scale wave height variations, in spite of its rather low rate for the nadir beam (5 Hz instead of 20 Hz for Jason), and the absence of Doppler processing.

⁴⁹² Appendix A Non-homogeneous H_s and waveform retracking

In this analysis we keep the most simple model of altimeter measurement that is 493 also used in section 2: we neglect antenna pattern, thermal noise and and mispointing 494 effects, and neglect the Earth sphericity. These assumptions are meant to simplify the 495 algebra as much as possible while keeping the essential features of non-homogeneity in 496 wave heights. Likewise we have used the most simple cost function when fitting the wave-497 form, while maximum likelihood methods are generally used with real data (Rodriguez, 498 1988; Halimi, 2013). We also start by ignoring speckle noise. The analysis performed be-499 low is easily extended to consider the third parameter which is usually estimated in re-500 tracking wave forms, that is the Normalized Radar Cross Section. 501

A1 Wave groups and H_s estimate

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We consider a small perturbation Δ_H of H_s over an area A, localized around a range $h+R_0$. The original normalized Brown waveform of eq. (7) corresponds to the histogram of the ocean area per unit range, divided by $2\pi h$ so that it varies between 0 and 1, with h the satellite altitude. Thus the perturbation to the waveform is equivalent to removing the original Gaussian distributed with $\sigma_H = H_s/4$ and replacing it by a new Gaussian with $\sigma' = (H_s + \Delta_H)/4$, multiplied by the area A and divided by the normalization factor $2\pi h$. We define the parameter $a = A\Delta_H/(8\pi h)$, which should be small compared to H_s^2 . For a small change in H_s , the change in waveform is proportional to the derivative of the Gaussian distribution with respect to σ_H and we find that the waveform is now

$$w(R) = w_B(R, \sigma_H) + a \frac{\mathrm{e}^{-(R-h-R_0)^2/(2\sigma_H^2)}}{\sqrt{2\pi}} \frac{(R-h-R_0)^2 - \sigma_H^2}{\sigma_H^4} + O(a^2)$$
(A1)

We note that a smaller change Δ_H over a larger area A changes the waveform in the same way as a larger change over a smaller area, provided that a is the same. For simplicity we redefine the Chelton footprint diameter as $r'_C = \sqrt{2H_sh}$, and we find that taking an area of radius $\alpha r'_C$ gives $a = \alpha^2 \Delta_H H_s/4$.

The shape of these distorted waveforms is illustrated in Fig. A1. With the exaggerated distortion shown here, fitting a Brown waveform would give a wave height of $H_{s,\text{fit}} =$ 12.6 m for $R_0 = 2.5$ m and $H_{s,\text{fit}} = 10$ m for $R_0 = 0$, which is a strange way to average the $H_s = 13$ m over part of the footprint and 10 m in the rest of the footprint. For smaller values of the perturbation a, the deviation in the fitted H_s can be computed analytically.

For simplicity we will assume that the waveforms are defined for $-\infty < R < \infty$, and the sum of the difference squared between w(R) and $w_B(R, \sigma'_H)$, when integrated



Figure A1. Example waveforms in the presence of a localized change in H_s around the range $h + R_0$, for $H_s = 10$ m. In order to be visible, the perturbation was exaggerated taking $a = 1.875 \text{ m}^2$, that would correspond to $\Delta_H=3$ m over an area of radius $r'_C/4$, a perturbation that is neither small nor localized.

from $R = -\infty$ to $R = -\infty$ is the following cost function,

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$$\begin{split} C &= \int_{-\infty}^{\infty} \left\{ [w_B(R,\sigma_H) - w_B(R,\sigma'_H)] + [w(R) - w_B(R,\sigma_H)] \right\}^2 \mathrm{d}R \\ &\simeq \int_{-\infty}^{\infty} \left\{ (\sigma_H - \sigma'_H) \frac{\partial w_B(R,\sigma_H)}{\partial \sigma_H} + [w(R) - w_B(R,\sigma)] \right\}^2 \mathrm{d}R \\ &= (\sigma'_H - \sigma_H)^2 \frac{1}{4\sqrt{\pi}\sigma_H} + (\sigma'_H - \sigma_H) \frac{aR_0}{8\sqrt{\pi}\sigma_H^5} \mathrm{e}^{-R_0^2/(4\sigma_H^2)} (R_0^2 - 6\sigma_H^2) + \frac{3a^2}{8\sqrt{\pi}\sigma_H^3} \mathrm{e}^{-R_0^2/(4\sigma_H^2)} \mathrm{e}^{-R_0^2/(4\sigma_H^2$$

Fitting σ'_H corresponds to solving $\partial C/\partial(\sigma'_H - \sigma_H) = 0$. We note that that error terms that are either not a function of $(\sigma'_H - \sigma_H)$ or odd functions of R have no impact on the fitted value. For example the a^2 term in eq. (A1) does not contribute any difference to the fit.

We find that the fitted value differs from the background value H_s by a factor proportional to a and function of R_0/H_s ,

$$H_{s,\text{fit}} = H_s + \frac{A}{\pi h} \frac{\Delta_H}{H_s} \underbrace{\left[2\frac{R_0}{H_s} \left(6 - \left(\frac{4R_0}{H_s}\right)^2 \right) e^{-4R_0^2/H_s^2} \right]}_{J(R_0/H_s)},\tag{A2}$$

with the function J in brackets having a maximum close to 2 for $R_0 \simeq H_s/4$, as shown in Fig. A2.

We note that this perturbation is zero for $R_0 = 0$, meaning that a localized change at the center of the footprint¹ does not modify the estimated H_s . This lack of impact

 $^{^{1}}$ In a similar way, but for different reasons, seismic waves travelling through a heterogeneous medium are most sensitive to variations in speed not exactly in the middle the ray path but it at a fraction of the



Figure A2. Functions $J(R_0/H_s)$ and $J_2(R_0/H_s)$ corresponding to the term in square brackets in eqs. (A2) and (A6). The maximum of J is at $R_0/H_s = 0.5\sqrt{0.5(3-\sqrt{6})} \simeq 0.26$, where Jtakes a value close to 1.96. This location corresponds to a distance from nadir approximately $\sqrt{0.26}r'_C \simeq r'_C/2$.

on $H_{s,\text{fit}}$ comes from the fact that the perturbation of the waveform (the second term in eq. A1) is an odd function of range and thus orthogonal to the even functions that are the Brown waveforms with zero epoch $w_B(R, \sigma_H)$. The maximum perturbation of $H_{s,\text{fit}}$ occurs for H_s perturbations at a range R_0 close to σ_H , i.e. corresponding to a distance from nadir of $r'_C/4$. Eq. (A2) gives results that are fairly robust for finite values of a/H_s^2 , and would predict a wave height of 12.9 m in the case $R_0 = 2.5$ m shown in Fig. A1.

We now consider the average effect of the perturbation by computing the average over R_0 , taking all values of R_0 from 0 to nH_s , which corresponds to averaging over an area $B = \pi n r'_C{}^2 = 2n\pi H_s h$. The integral of the function in brackets is

$$I = \int_0^\infty 2\frac{R_0}{H_s} \left(6 - 16\frac{R_0^2}{H_s^2} \right) e^{-4R_0^2/H_s^2} dR_0 = 0.5H_s.$$
(A3)

As a result, the average effect of a Δ_H change over an area $A = \pi \alpha^2 r'_C{}^2 = 2\pi \alpha^2 h H_s$ is, when *n* is large,

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$$\delta_{H,\text{alti}} = \frac{1}{nH_s} \int_0^{nH_s} \left(H_{s,\text{fit}} - H_s \right) \mathrm{d}R_0 = \frac{1}{2n} \frac{A}{\pi h} \frac{\Delta_H}{H_s} = \frac{\alpha^2}{n} \Delta_H. \tag{A4}$$

This average effect of the localized perturbation of H_s is the same as a true area average, which is the perturbation times the ratio of the areas A and B, namely $\delta_H = \Delta_H A/B$. In other words, the perturbation is amplified if located at $0.15 < r/r_C' < 0.34$ from nadir, by a factor J that is up to 2. Otherwise the perturbation is attenuated, so that on average it is equal to the true perturbation. The unbiased estimate of $H_{s,fit}$, with a perturbation that changes sign when Δ_H changes sign, and this averaging property are specific to the simple least squares used here. For example, fitting the logarithm of the

wavelength from the ray (Marquering & Nolet, 1999). Somehow the "range-blurring" associated with the wave displacement in altimetry is similar to the finite-frequency effect in seismology.

waveform produces a biased estimator and a non-zero response for $R_0 = 0$. Hence the results presented here are specific to the fitting method.

In practice, distributed anomalies of H_s are not only a function of the distance from nadir, so that a local estimate of H_s will combine positive and negative anomalies Δ_H that are located at the same distance from nadir, and will partially cancel. This explain that our best fit for r_0 is $r_C/4.5$, smaller than the $r_C/2$ which is a more typical scale of the footprint. A spatial filter that would behave like the altimeter retracker can be built from the J function, converting the range $h + R_0$ to a horizontal distance from nadir $r = \sqrt{2hR_0}$.

A2 Wave groups and sea level estimate

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⁵⁵⁸ While perturbations at nadir do not change the H_s estimate, they would change ⁵⁵⁹ the mean sea level z_e (the epoch is $-z_e$) when using a 2-parameter waveform

$$w_{B2}(R,\sigma_H, z_e) = \frac{1}{2} \left[1 + \operatorname{erf}\left(\frac{(R+z_e) - h}{\sqrt{2}\sigma_H}\right) \right].$$
(A5)

In the case shown in Fig. A1 with $R_0 = 0$ the estimated mean sea level is z = -37 cm. 561 We thus expect wave groups to contribute to fluctuations in the estimated sea level at 562 the scale of groups. The estimation of that effect follows the same method used above. 563 Fitting $w_{B2}(R, \sigma'_H, z_e)$ to our waveform w(R) given by eq. (A1) is obtained by minimiz-564 ing a modified cost function, that is the same as C but with one extra term $z_e \partial w_{B2} / \partial z_e$ 565 inside the curly brackets, giving two extra non-zero terms proportional to z_e^2 and z_e . We 566 note that the cross-term proportional to $z_e(\sigma'_H - \sigma_H)$ is an odd function of R and thus 567 integrates to zero. After integration over R we get the cost function, 568

$$C_{2} = C + \frac{z_{e}^{2}}{2\sqrt{\pi}\sigma_{H}} + \frac{az_{e}}{4\sqrt{\pi}\sigma_{H}^{2}} e^{-R_{0}^{2}/(4\sigma_{H}^{2})} \left(\frac{R^{2}}{\sigma_{H}^{2}} - 2\right).$$

Taking the derivative of C_2 with respect to z_e gives

$$z_e = -\frac{A\Delta_H}{2\pi hH_s} \underbrace{\left[\left(2 - 16\frac{R_0^2}{H_s^2} \right) e^{-4R_0^2/H_s^2} \right]}_{J_2(R_0/H_s)}.$$
 (A6)

The function J_2 is plotted in Fig. A2. Hence z_e has the strongest deviation when the wave height perturbation is centered at nadir, and the sign of the deviation is opposite to Δ_H : i.e. a wave group centered at the nadir would give a spurious lower sea level. On average the z_e deviation is zero mean when R_0 is varied. As a result of the different shapes of J and J_2 , there is no simple correlation of the H_s and z_e perturbations, contrary to the correlations induced by speckle noise in the waveform measurement (Sandwell & Smith, 2005).

There is some correlation for R_0/H_s between 0.7 and 1.2 which may contribute to 579 anti-correlation of sea level anomalies and wave heights at scales around r_{C} , and thus 580 may persist in 1 Hz data. We insist that these are spurious sea level variations. In deep 581 water these spurious oscillations are much larger than the fraction of a millimeter asso-582 ciated to true sea level variations with bound infragravity elevation that is anti-correlated 583 with the envelope of kilometer-scale wave groups (Ardhuin et al., 2004). The spurious 584 sea level oscillations described are also probably generally larger in amplitude than the 585 larger scale (20-km wavelength) true sea level variations associated to free infragravity 586 waves that have no phase correlation with the local envelope (Ardhuin et al., 2014). In 587 shallow water, the real sea level fluctuations can be more important. 588

589 A3 Speckle noise

Random fluctuations in the electromagnetic power measured by the radar combine an additive thermal noise that can often be neglected and a multiplicative noise that is caused by the Rayleigh fading of the interfering reflections off a random sea surface (Quartly et al., 2001). In fact speckle is to the radar power what wave groups are to the wave energy. A good model for the speckle is a multiplicative random noise, so that the measured waveform for each range is multiplied by a factor $(1+\varepsilon(R))$ with $\varepsilon(R)$ following a χ^2 distribution with N(R) degrees of freedom depending on the number of pulses averaged and the pulse repetition frequency (Quartly et al., 2001).

For the retracking, the effect of this speckle perturbation is one additional term $\varepsilon(R)w(R)$ inside the curly brackets of the cost function. Expanding the square and expressing the integral, it gives two terms, one proportional to $(\sigma'_H - \sigma_H)$ that is relevant to the H_s estimate and the other proportional to z_e fit, so that the cost function is now,

$$C_{3} \simeq C_{2} - 2(\sigma_{H}' - \sigma_{H}) \int_{-\infty}^{\infty} \varepsilon(R) w(R) \frac{\partial w_{B2}}{\partial \sigma_{H}} dR - 2z_{e} \int_{-\infty}^{\infty} \varepsilon(R) w(R) \frac{\partial w_{B2}}{\partial z_{e}} dR, \quad (A7)$$

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$$\frac{\partial w_{B2}}{\partial \sigma_H} = -\frac{R-h+z_e}{\sigma_H^2 \sqrt{2\pi}} e^{-(R-h+z_e)^2/(2\sigma_H^2)},\tag{A8}$$

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$$\frac{\partial w_{B2}}{\partial z_e} = \frac{1}{\sigma_H \sqrt{2\pi}} e^{-(R-h+z_e)^2/(2\sigma_H^2)}.$$
(A9)

The estimated wave height that gives $\partial C_3 / \partial (\sigma'_H - \sigma_H) = 0$ thus has an extra term induced by speckle noise,

$$H_{s,\text{fit}} = H_s + \frac{A}{\pi h} \frac{\Delta_H}{H_s} J(R_0/H_s) + 16\sqrt{2}H_s \int_{-\infty}^{\infty} \varepsilon(u) \left(1 + \text{erf}(2\sqrt{2}u)\right) u e^{-8u^2} du, \quad (A10)$$

with $u = (R - h + z_e)/H_s$. The speckle-induced perturbation of $H_{s,\text{fit}}$ is a weighted 610 sum of random fluctuations with zero mean. In practice we can consider $\varepsilon(R)$ to be Gaus-611 sian, and the variance of the speckle perturbation is the sum of the variances associated 612 to each range R times the weigh squared. To get some useful order of magnitude we may 613 take the variance of $\varepsilon(R)$, which is 1/N(R), to be constant at 1/N. For large values of 614 H_s , the discretized waveform is well approximated by the continuous form and the part 615 of the variance of $H_{s,\text{fit}}$ induced by the speckle is approximately 5.0 H_s/N , with a stan-616 dard deviation $2.24\sqrt{H_s/N}$. Using the value N = 512 for the number of pulses of the 617 SWIM nadir beam that we may assume to be independent, and $H_s = 2$ m, this gives 618 a standard deviation of 0.14 m, broadly consistent with the background level in Fig. 8. 619 However, we note that the magnitude of the variability of $H_{s, \text{fit}}$ will depend on the method 620 used to fit the waveform. In the case of the SWIM data, the adaptive method that is used 621 is based on a maximum likelihood (Tourain et al., 2021). It is probably more robust to 622 speckle noise perturbations than the least square estimate used here, in particular for 623 this instrument that has a relatively high signal to noise ratio. 624

625 Acknowledgments

This research was made possible by a post-doctoral grant from CNES, and support from

ESA as part of the Sea State CCI project. CFOSAT data was provided by CNES and Ifremer.

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